



# IceCube - DeepCore - PINGU

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Penn State

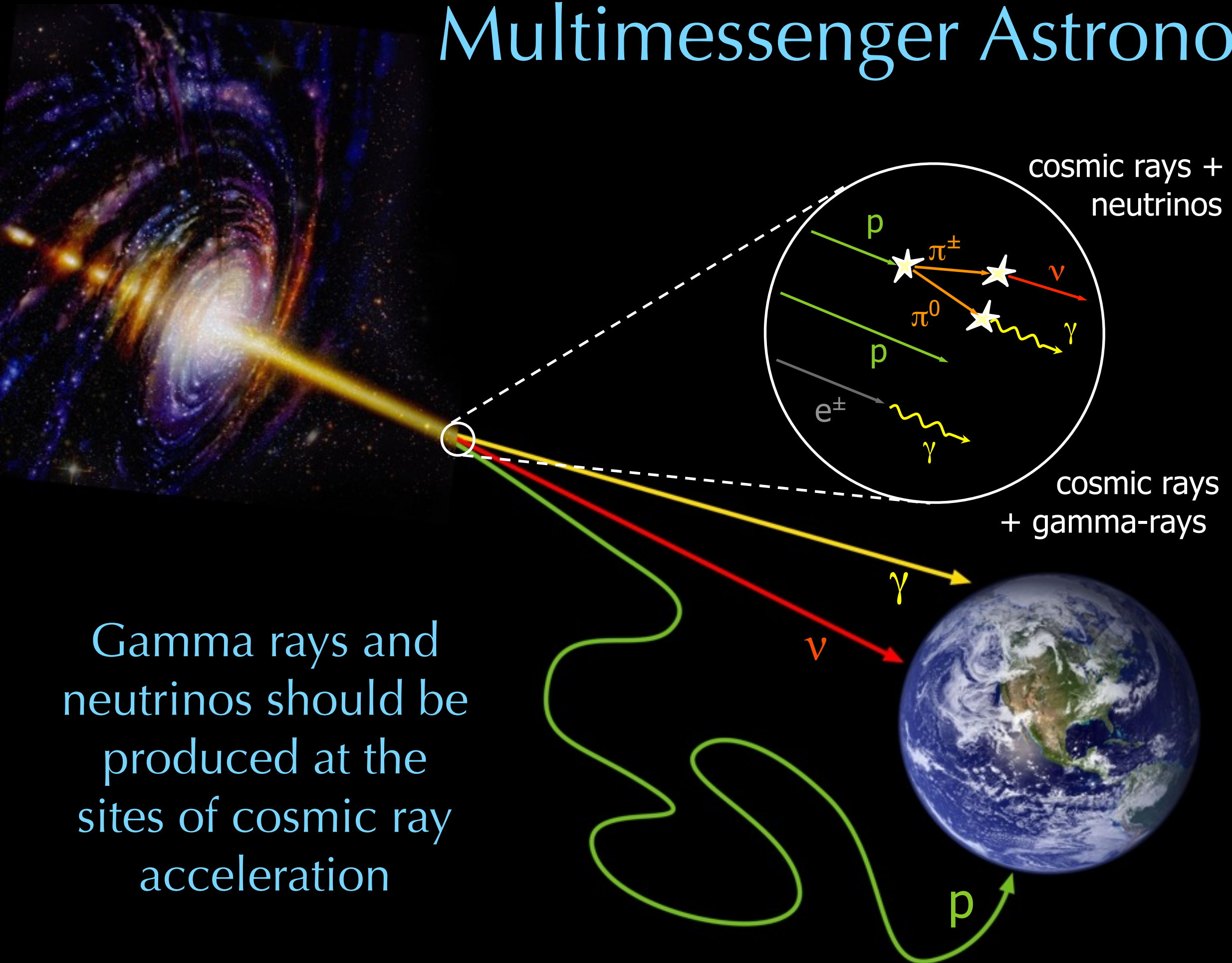
Implication of Neutrino Flavor Oscillations (INFO) 11  
July 2011

- IceCube astrophysical neutrino physics is alive and well
- DeepCore is an IceCube infill deployed to enhance sensitivity to neutrinos from  $\sim 10$  GeV to  $\sim 300$  GeV
  - Dark Matter
  - Neutrino Oscillations
- DeepCore is a multi-megaton scale neutrino detector at tens of GeV, which is situated inside a gigaton sized “veto”
- Proposed phased extensions
  - Phase 1 (PINGU) - Down to  $\sim 1$  GeV
  - Phase 2 (BeyondDC) - Down to  $\sim O(10)$  MeV

# IceCube Classic

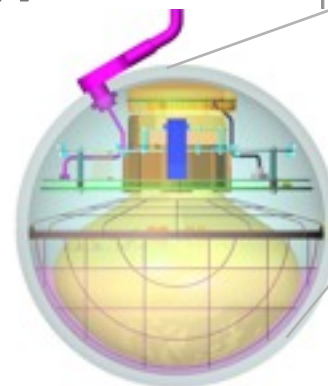
- Neutrinos are long distance cosmic messengers
  - Photons interact with CMB
  - Charged cosmic rays lose directionality through magnetic deflection
- The cosmic laboratory
  - Complementarity with cosmic rays
  - Astrophysical objects and Cosmic Ray acceleration, leptonic or baryonic? (GRB, SNR, PWN)

# Multimessenger Astronomy

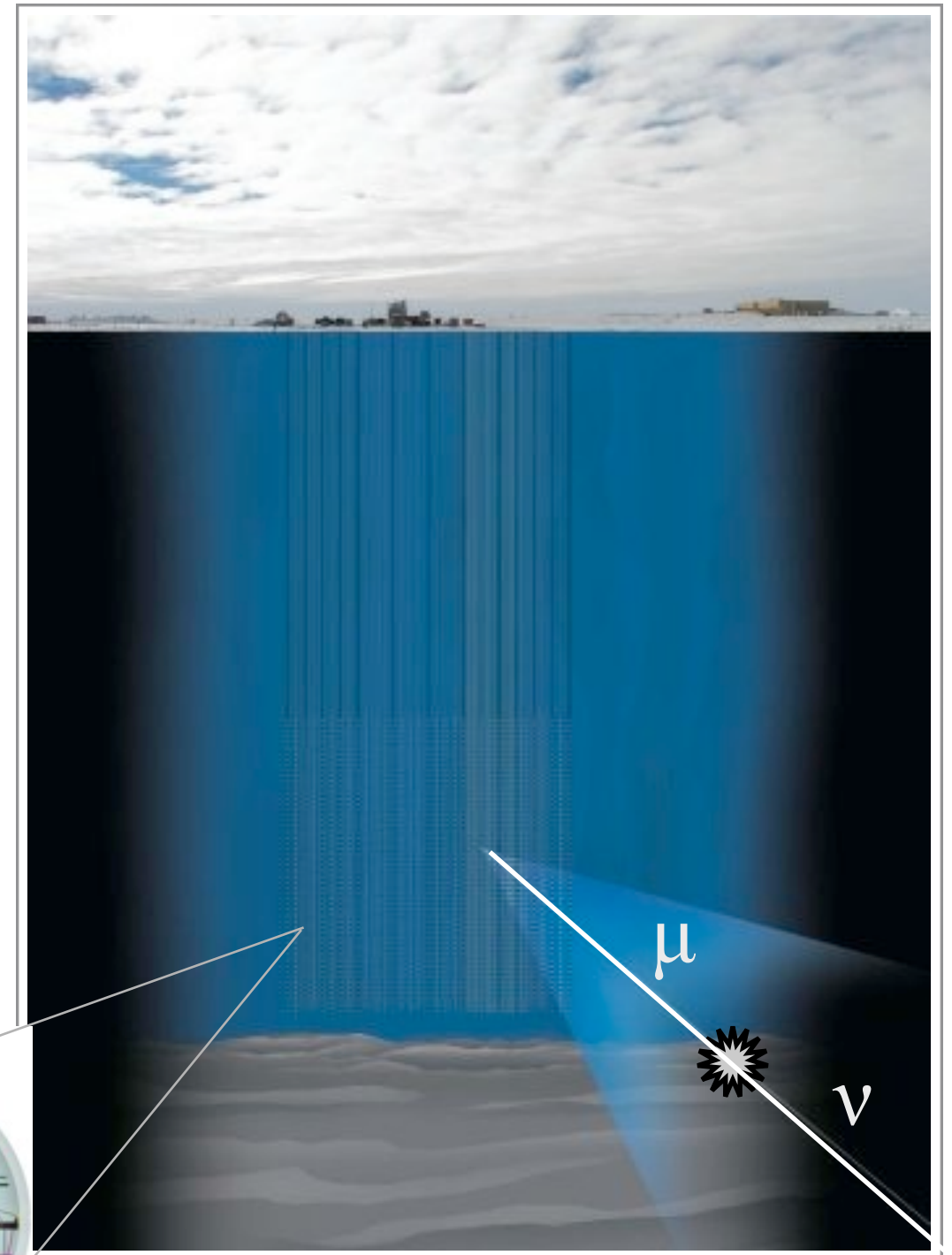


Gamma rays and neutrinos should be produced at the sites of cosmic ray acceleration

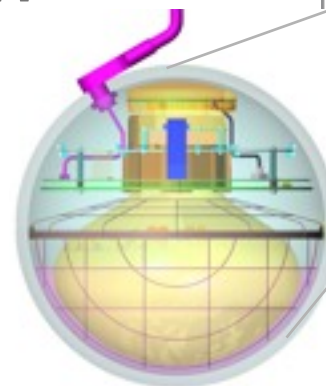
- $\sim 1\text{km}^3$  of instrumented ice
- Uses 5160 Digital Optical Modules (DOMs) across 86 strings within the ice to detect Cherenkov radiation
- 160 Cherenkov tank surface array (IceTop)
- Completed Dec. 18, 2010
- Deployed 1.5-2.5km below the surface



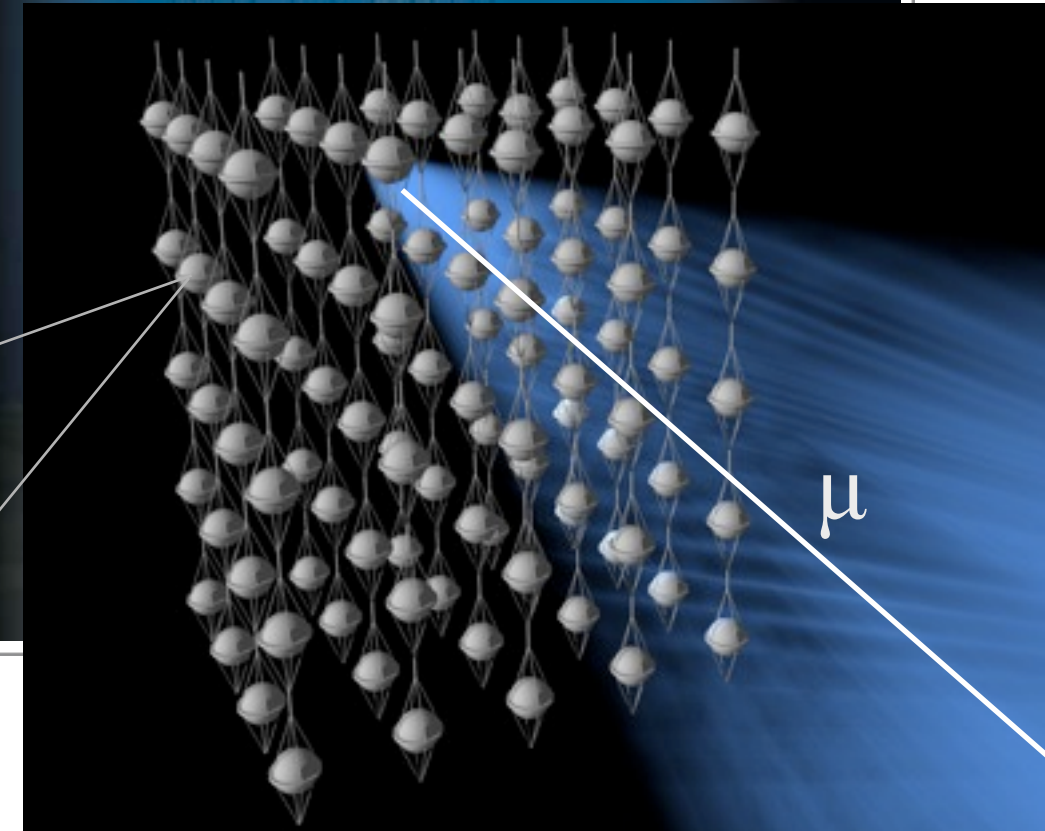
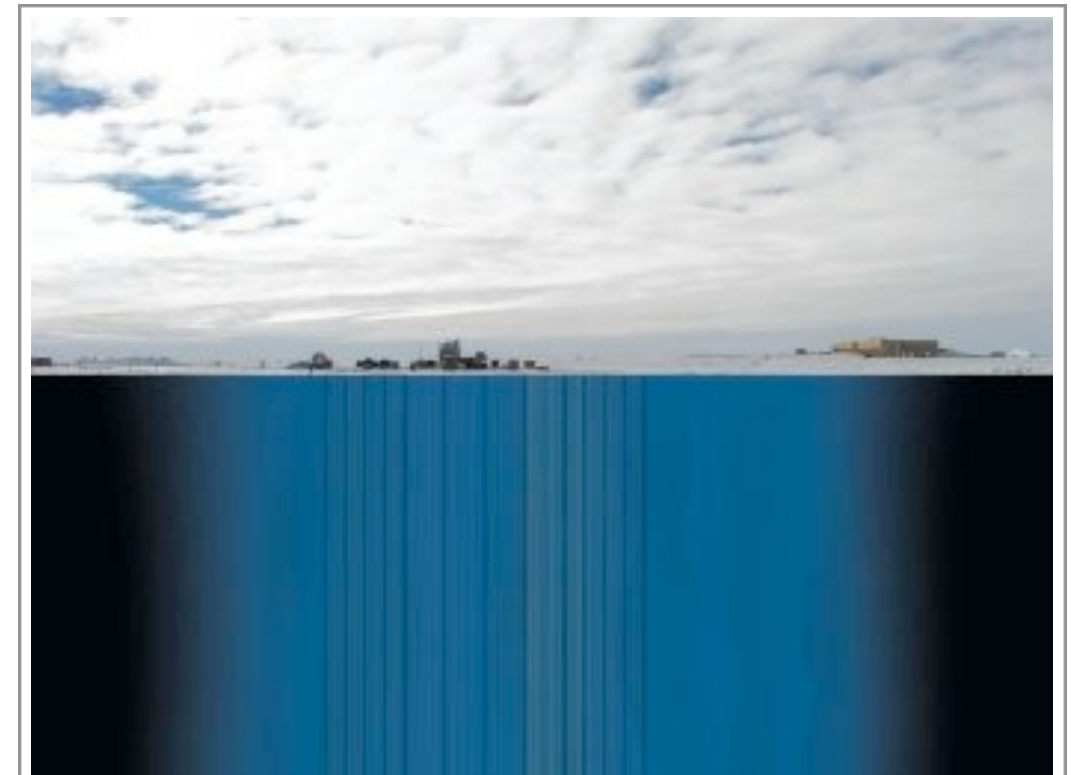
IceCube DOM

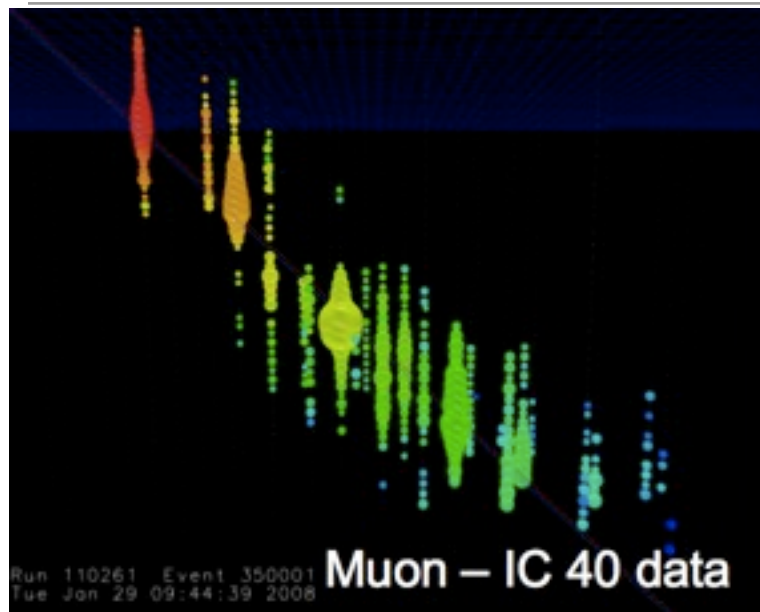


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IceCube DOM



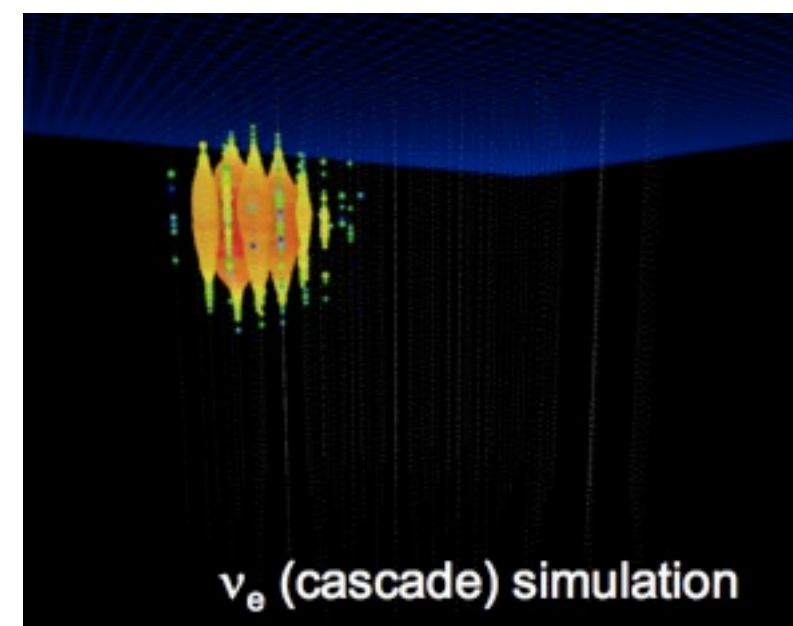


## Tracks:

- through-going muons

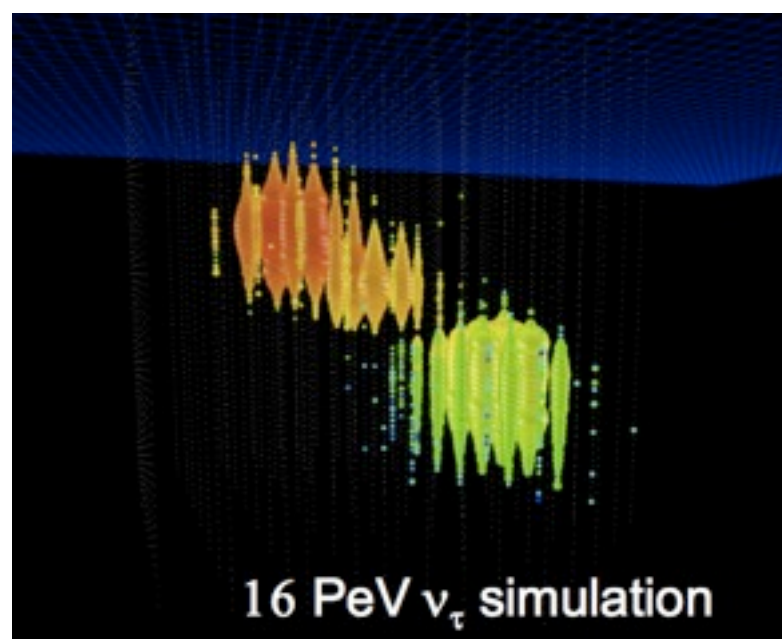
## Cascades:

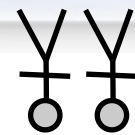
- Neutral current for all flavors
- Charged current for  $\nu_e$  and low-E  $\nu_\tau$



## Composites:

- Starting tracks
- high-E (PeV)  $\nu_\tau$  (Double Bangs)
- Good directional and energy resolution



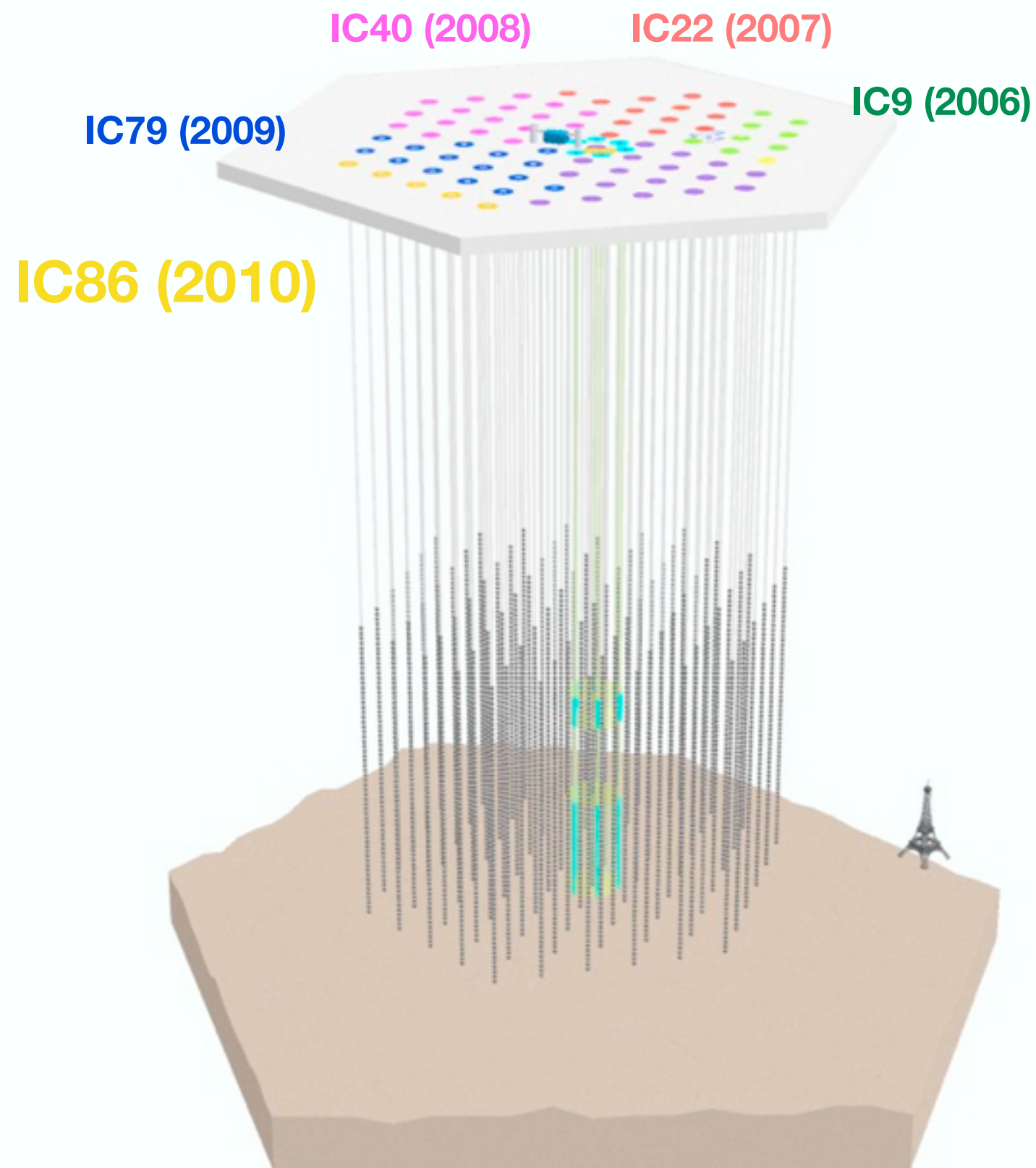


## The IceCube Collaboration

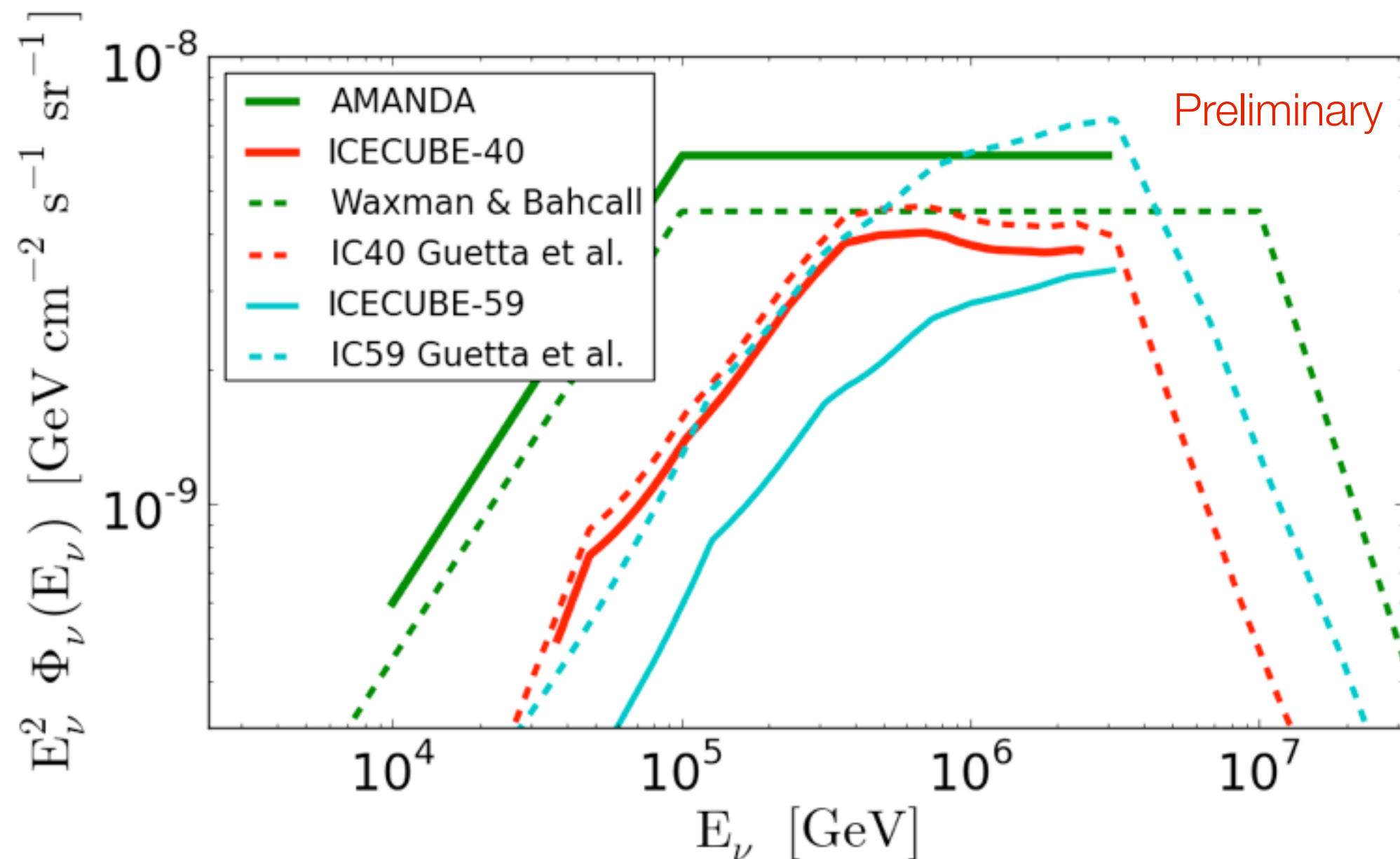
36 institutions - 4 continents - ~250 Physicists

# Data Sets by year

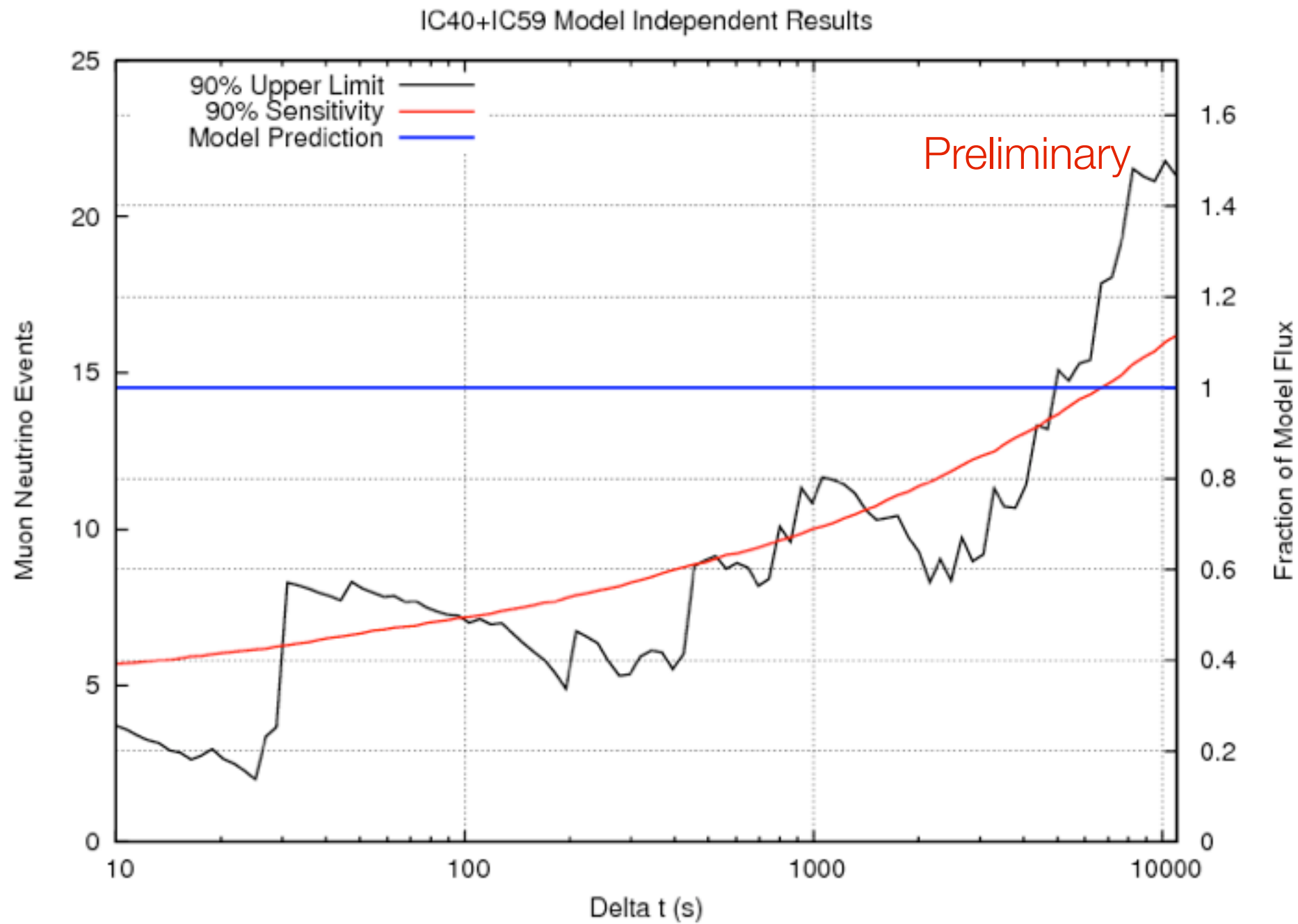
- IceCube
- DeepCore
- Beyond DeepCore



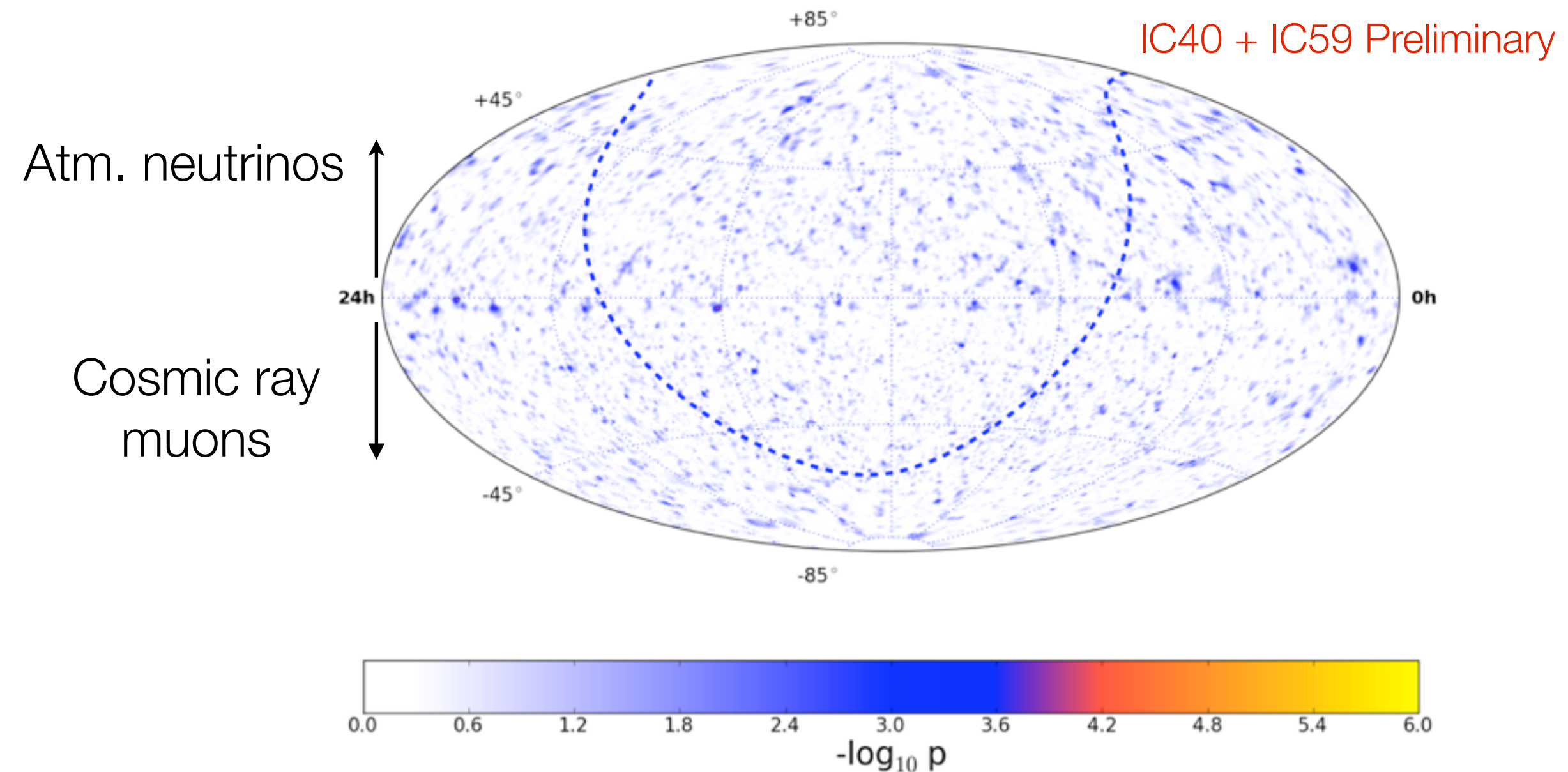
- Searches for neutrinos produced by  $p+\gamma$  interactions during the primary fireball



IC40 result  
arXiv:1101.1448



- Sky Map consistent with background



- IC59 Cosmic Ray Anisotropy - **arXiv:1105.2326**
- IC40 Diffuse Flux - **arXiv:1104.5187**
- IC40 Atmospheric Neutrino Spectrum - **arXiv:1010.3980v1**
- Supernova 2008D - **arXiv:1101.3942**
- IC22 Neutrino Induced Cascades - **arXiv:1101.1692**
- Neutrino Emission Constraints on 2010 Crab Flare - **arXiv:1106.3484**

# Below TeV+ Energies DeepCore



IceCube

# Below TeV+ Energies DeepCore



IceCube

# Below TeV+ Energies DeepCore



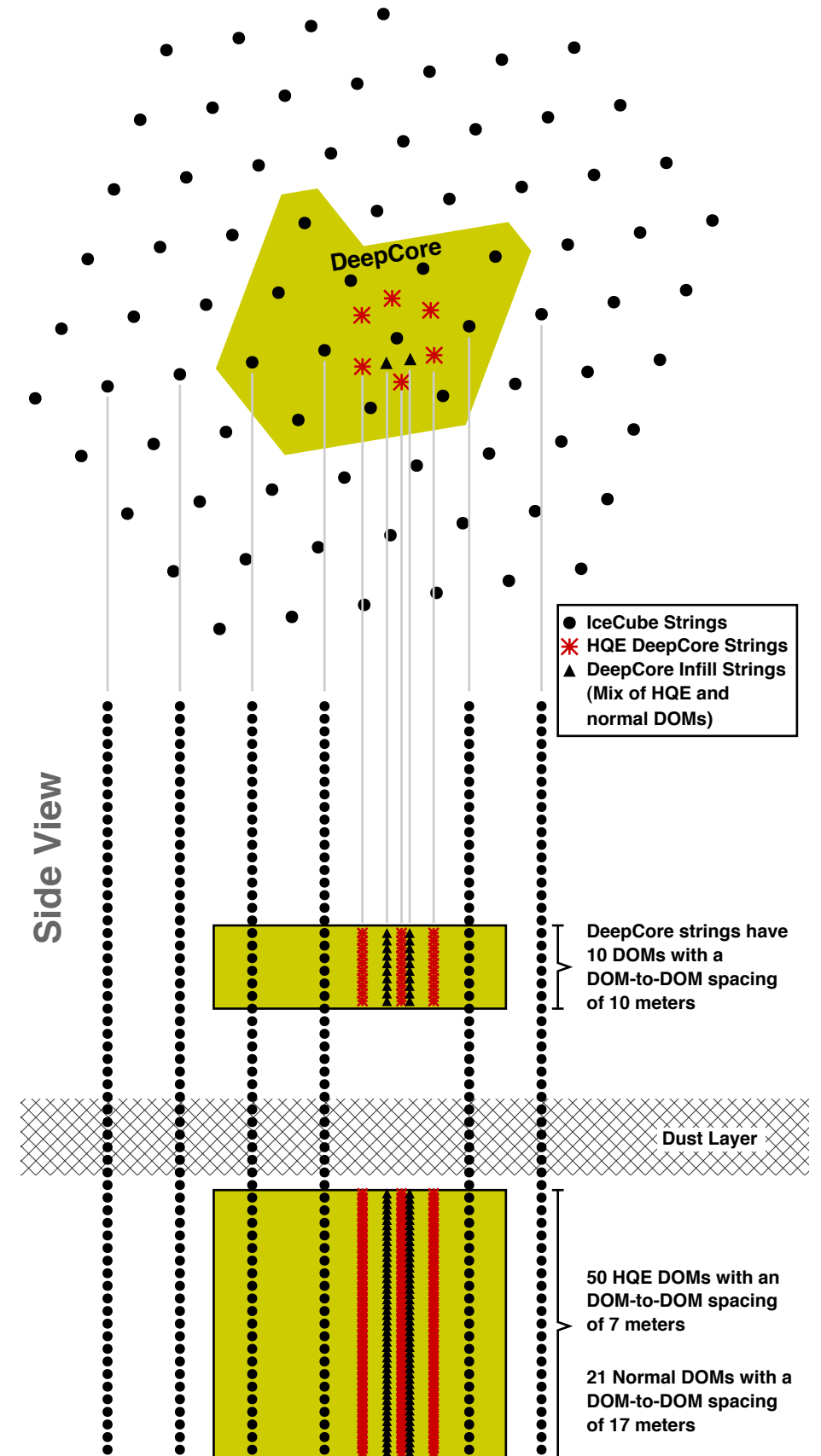
IceCube



DeepCore

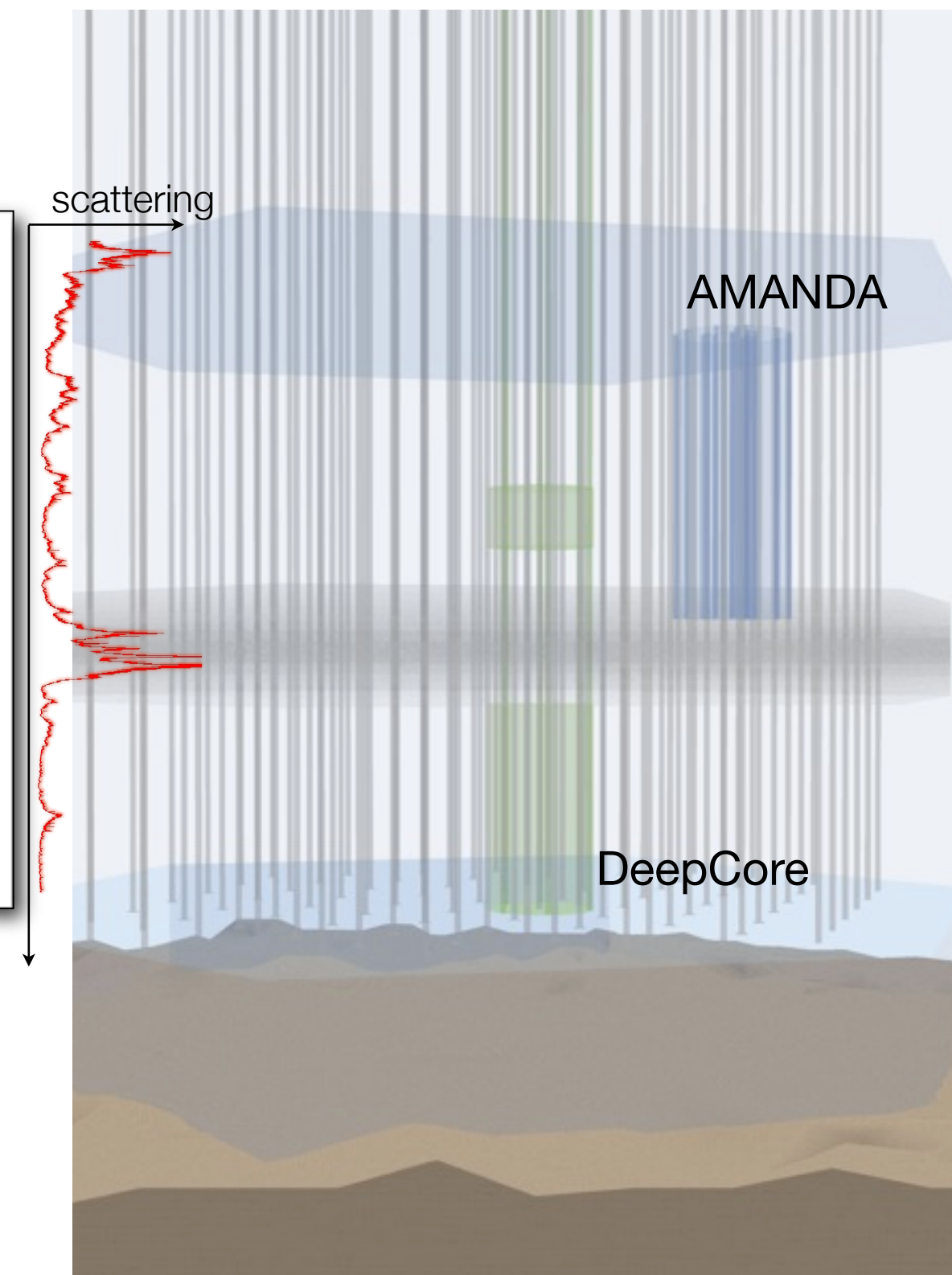
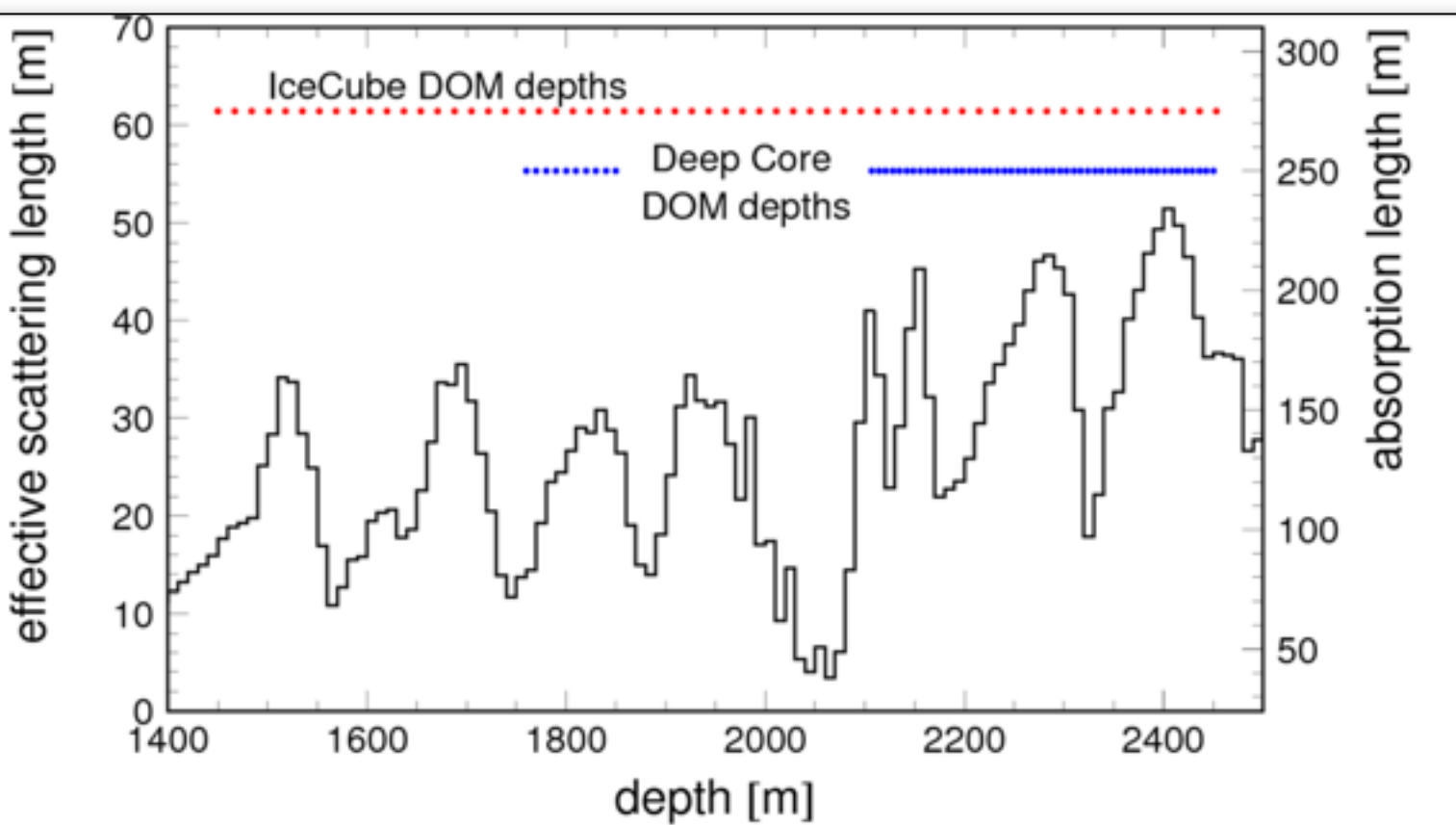
- 8 special Strings combined with 12 nearby standard IceCube Strings
- 72 m interstring spacing versus 125 m for IceCube
  - 7 m DOM spacing vs. 17 m
  - High Quantum Efficiency PMTs (35% higher QE)
- Deepest clearest ice
  - 40m scattering length
  - 140m absorption length

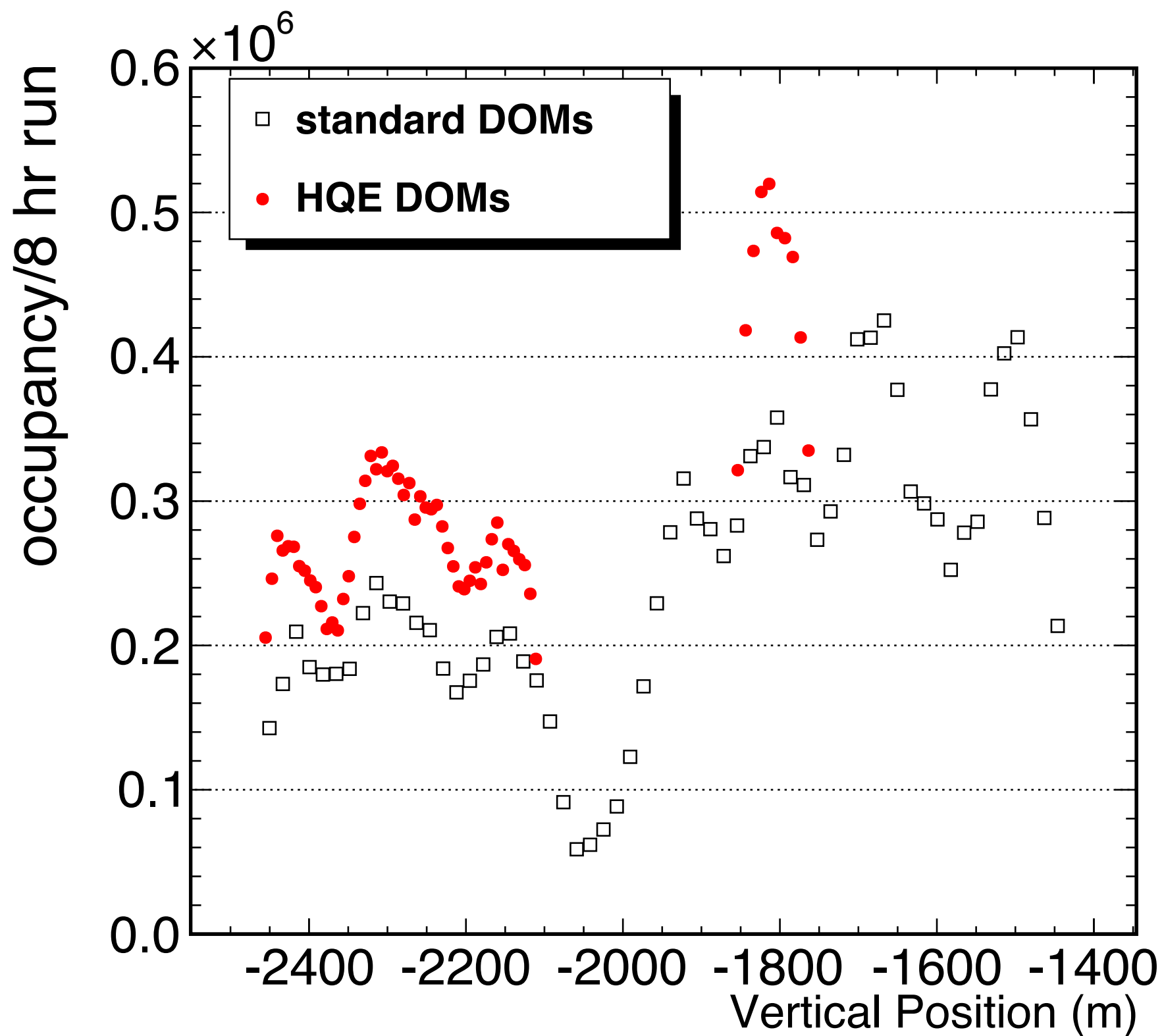
Overhead View



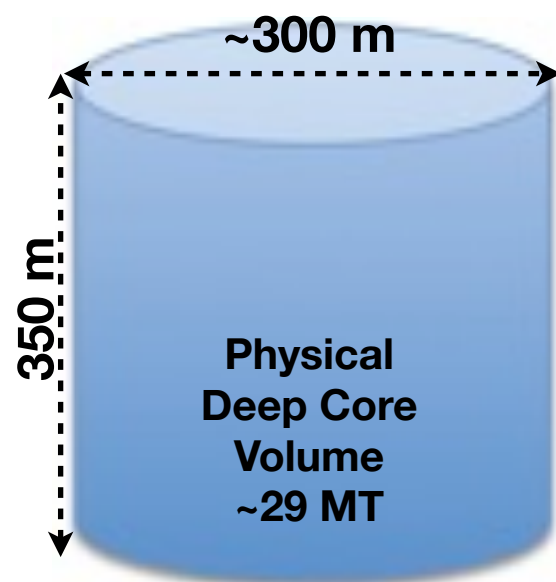
# Dust at Depth

- IceCube
- DeepCore
- Beyond DeepCore

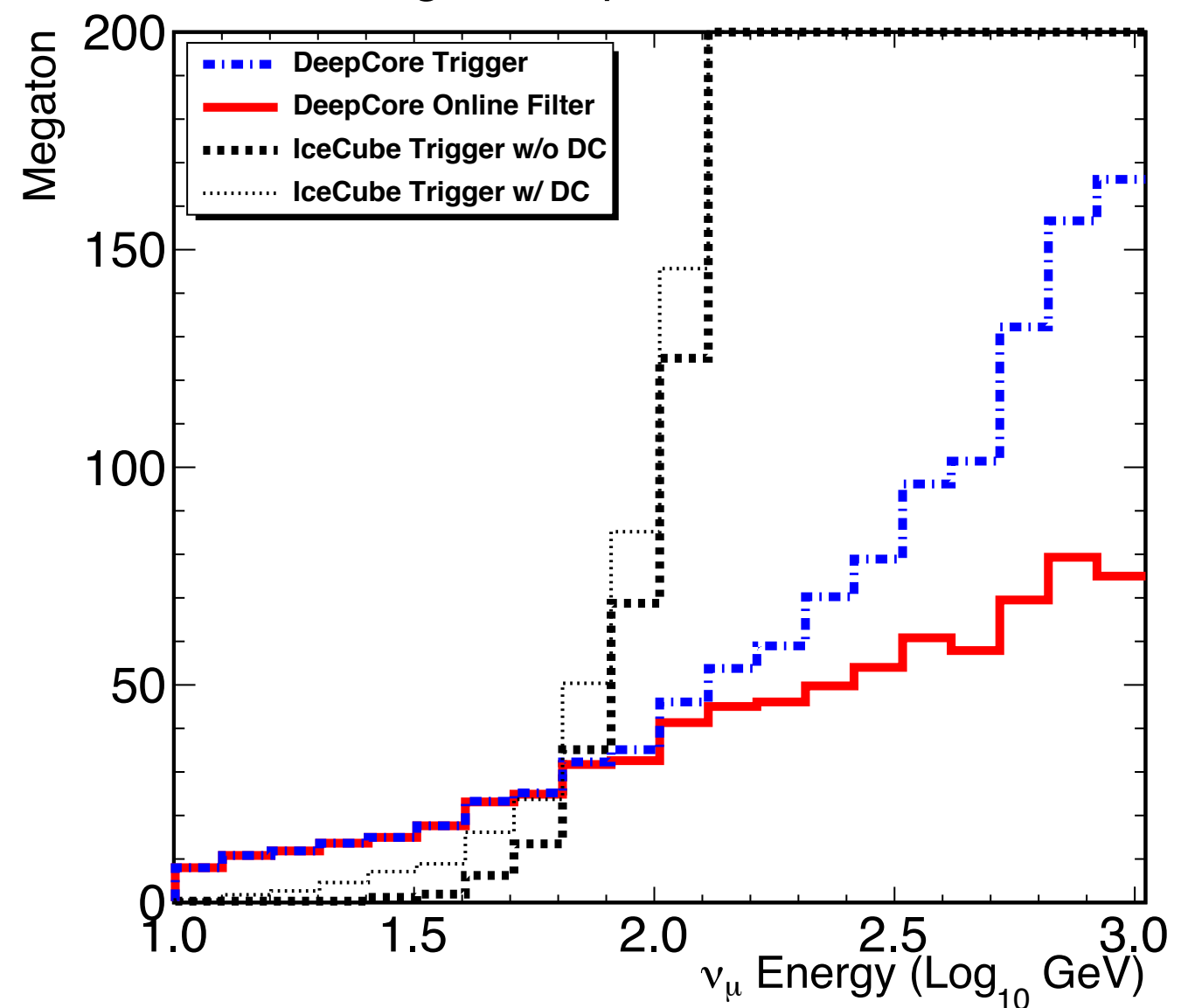




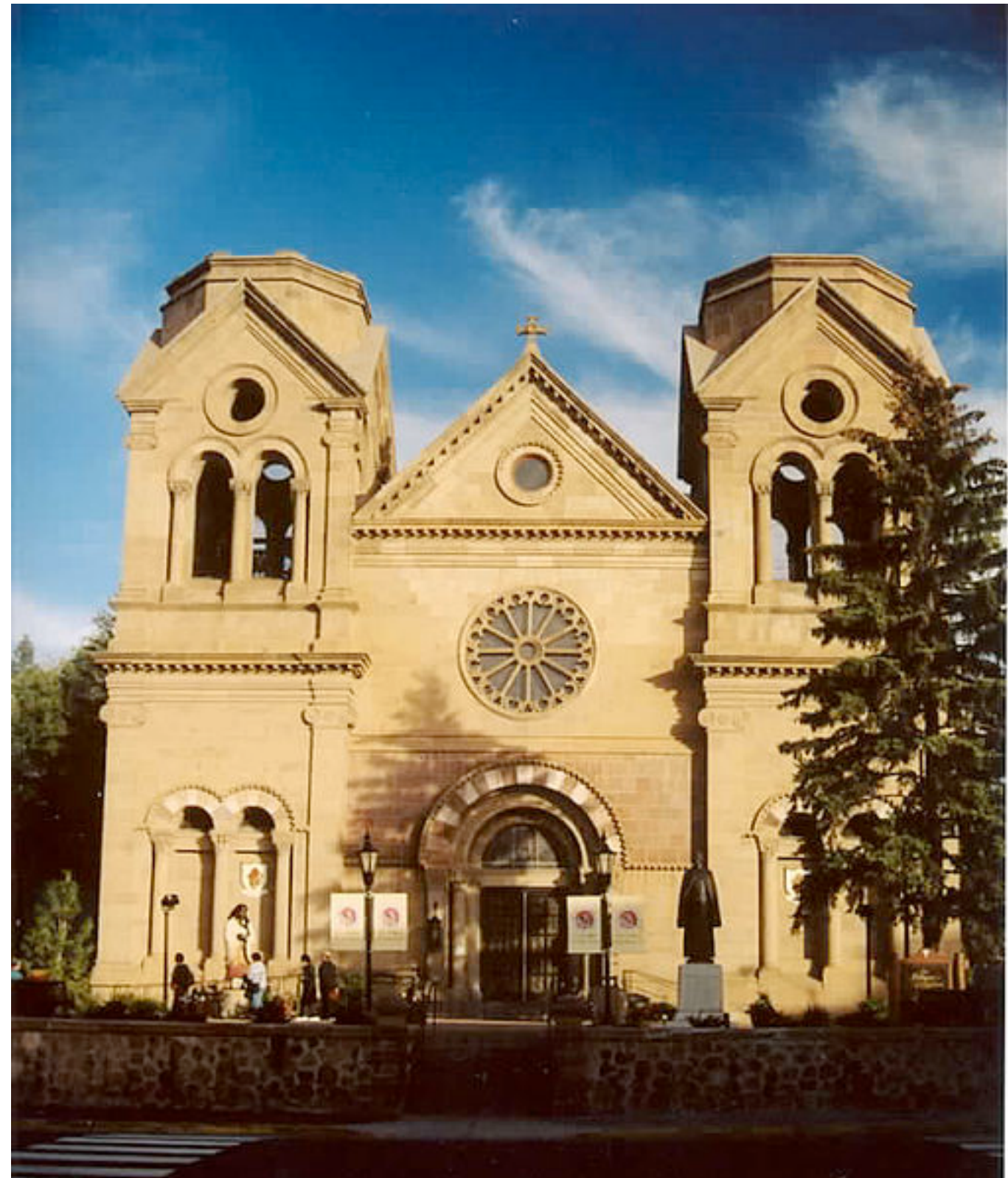
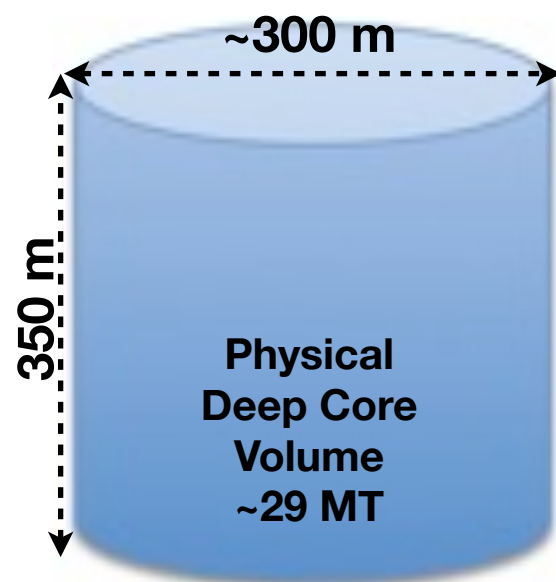
- Online Veto loosely constrains vertex to be within DeepCore volume
- Physical volume is ~29 MTon



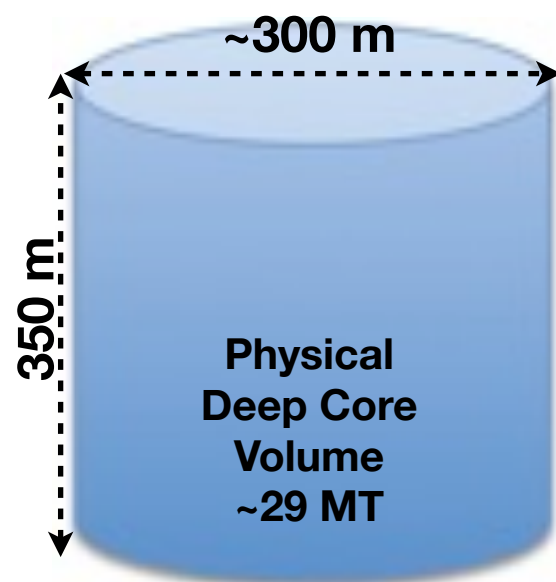
Effective volume for muons from  $\nu_\mu$  interacting in Deep Core



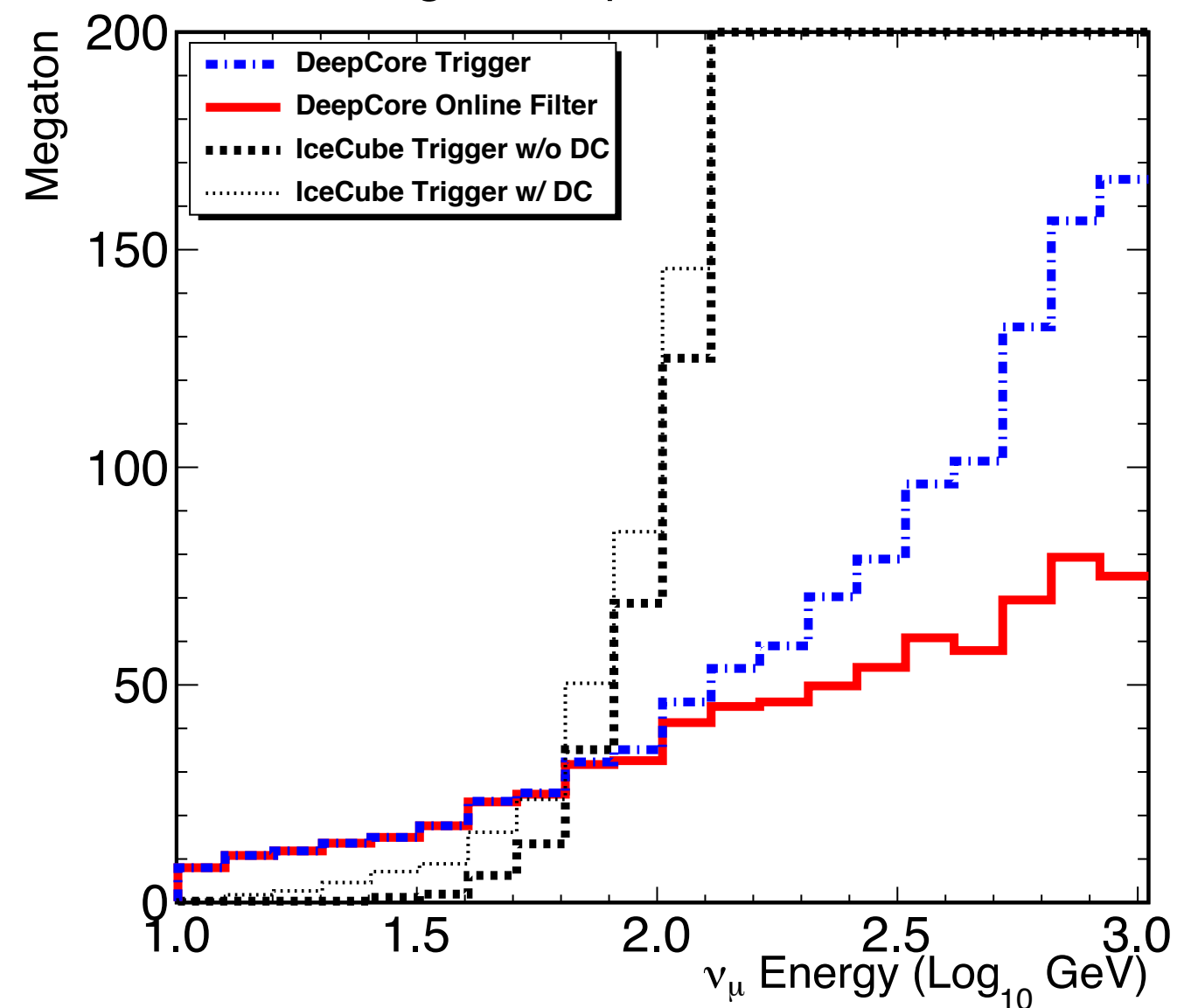
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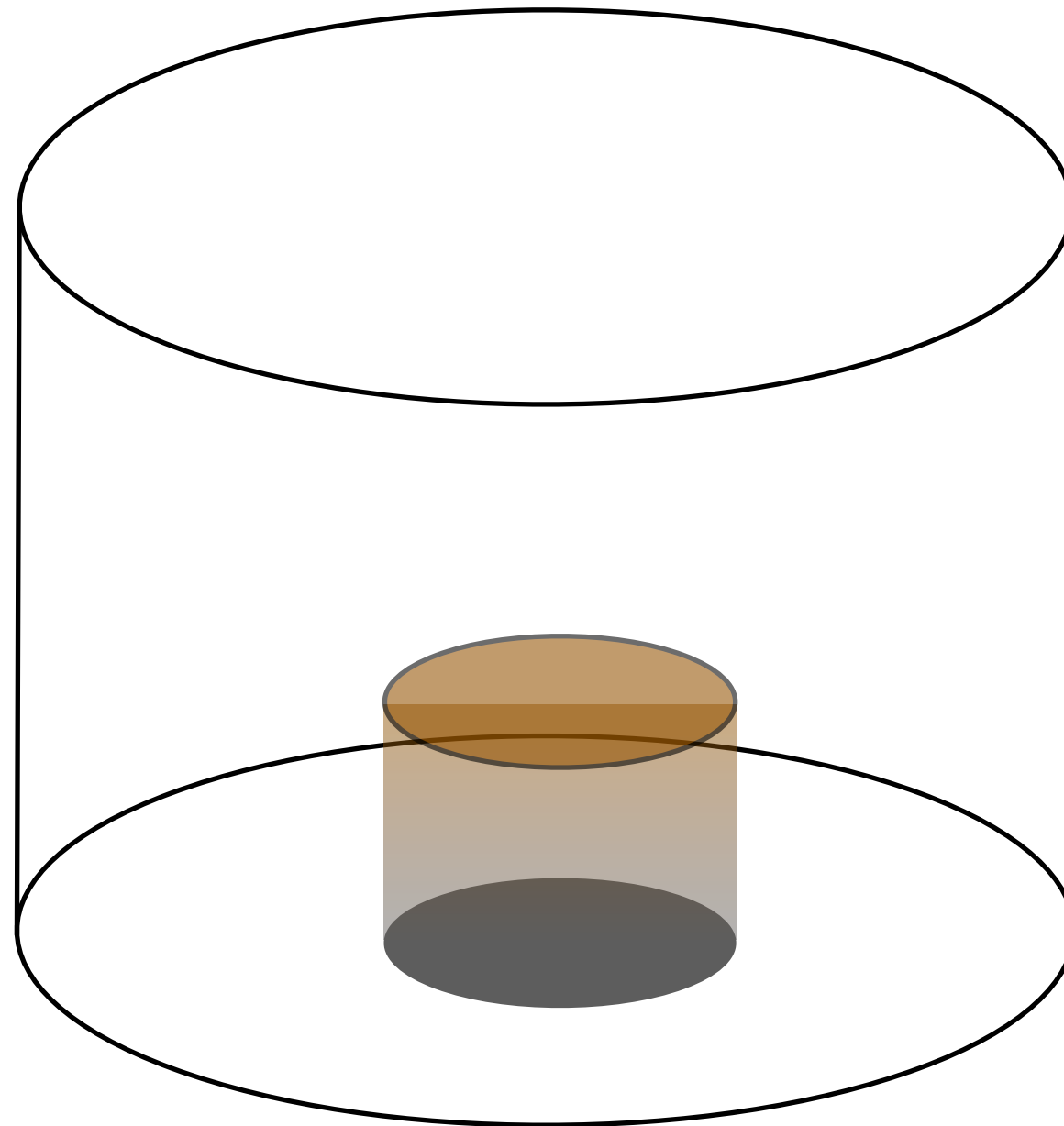


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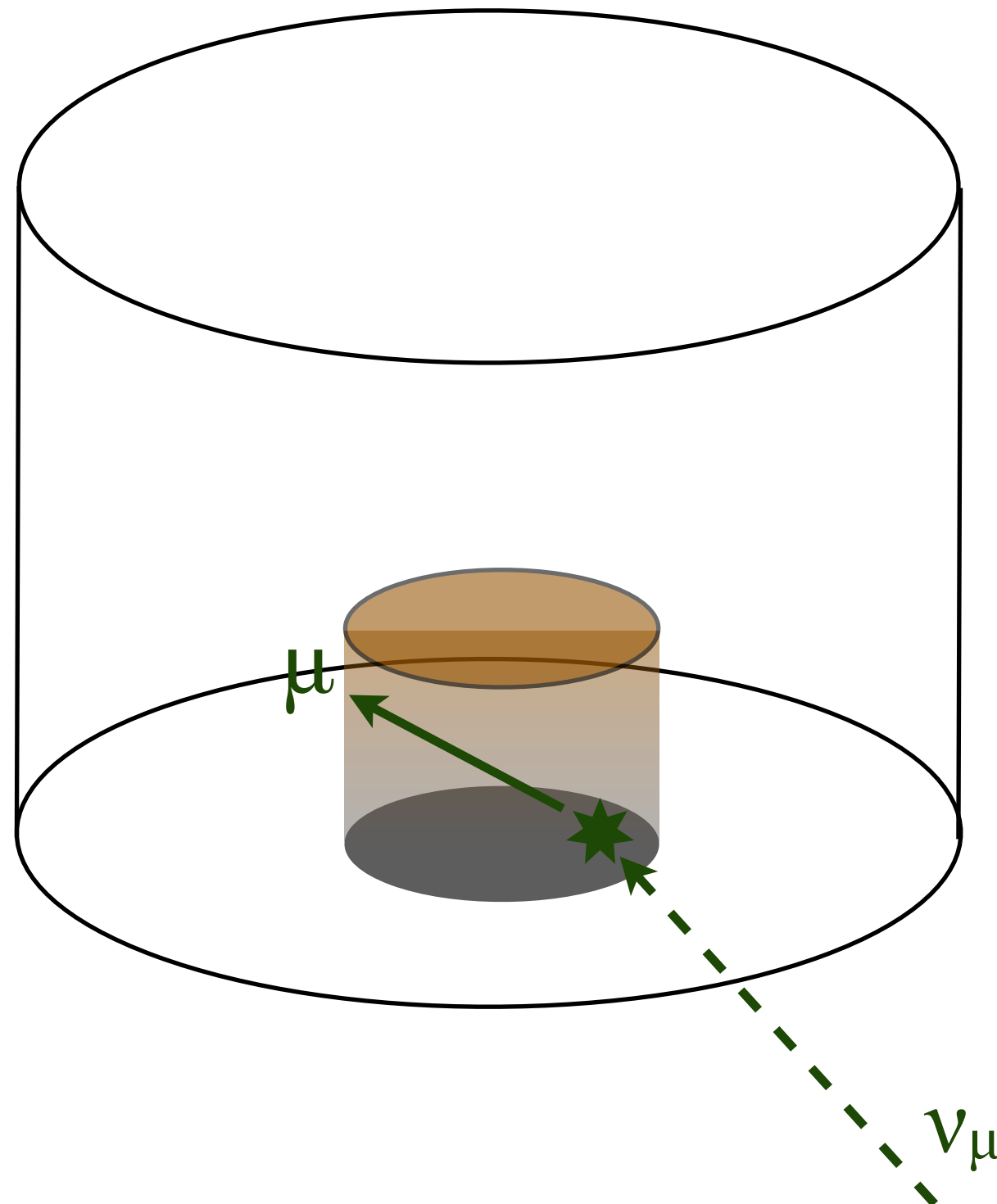
# More Signal = More Bkg

- IceCube
- DeepCore
- Beyond DeepCore



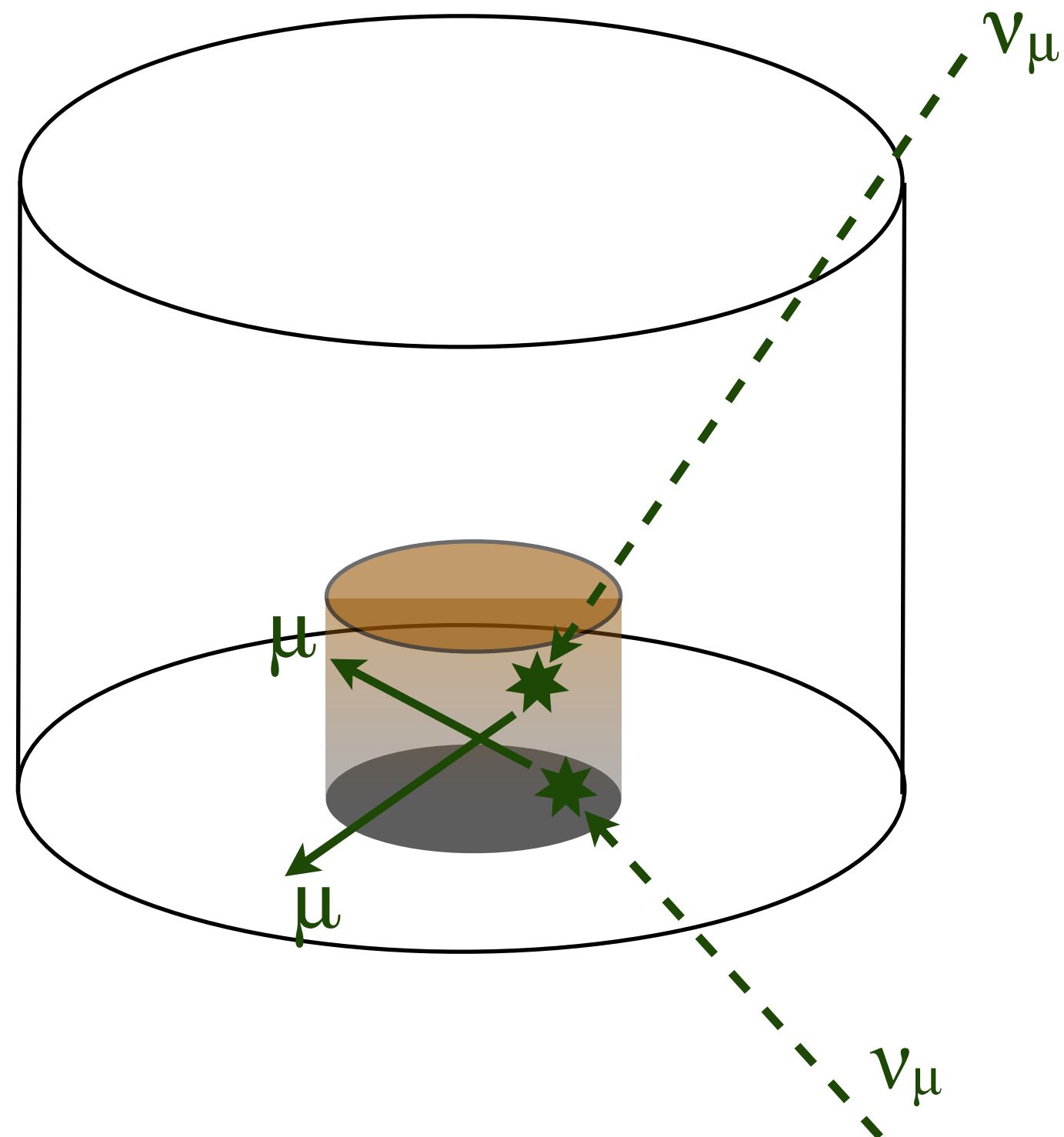
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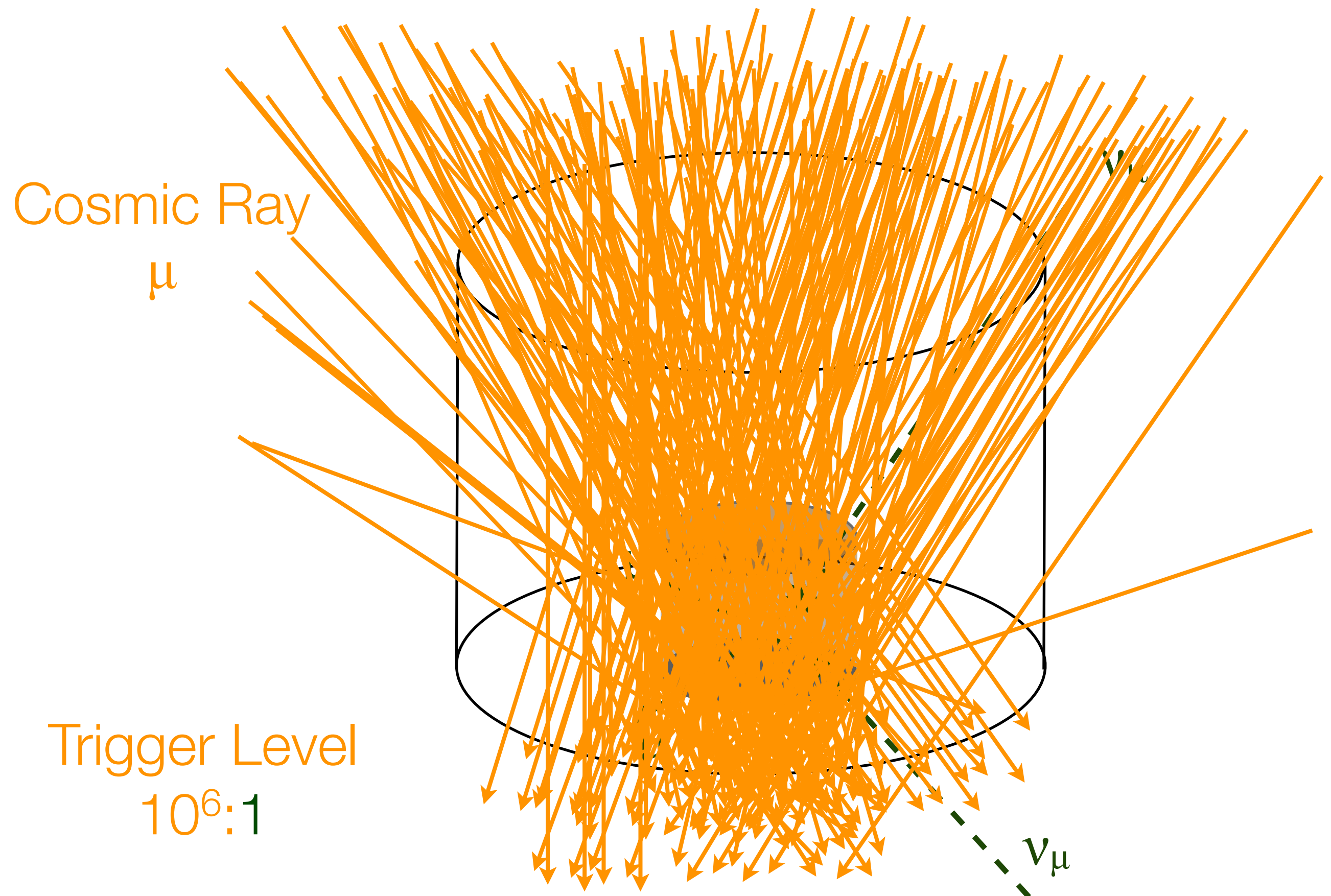
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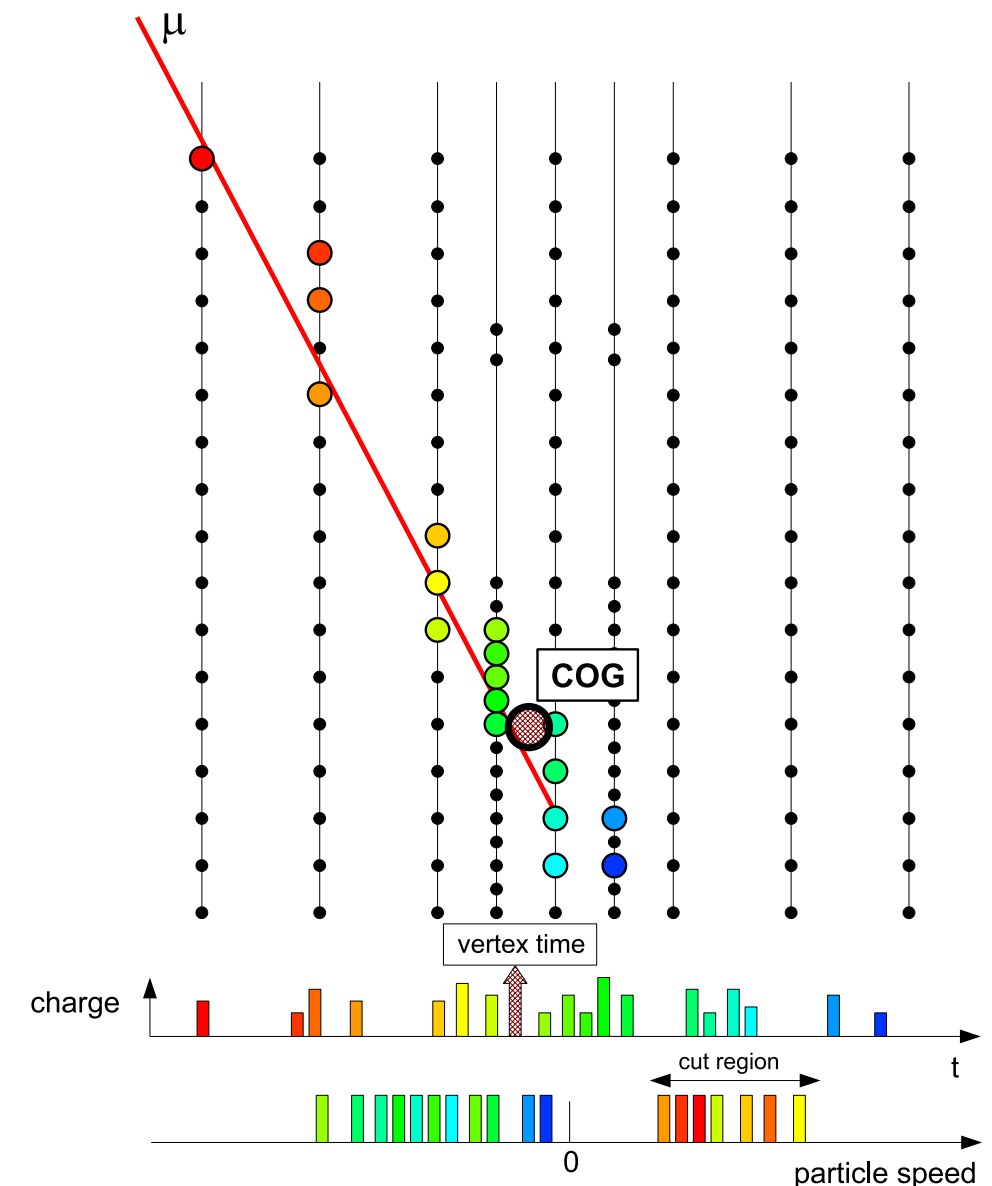


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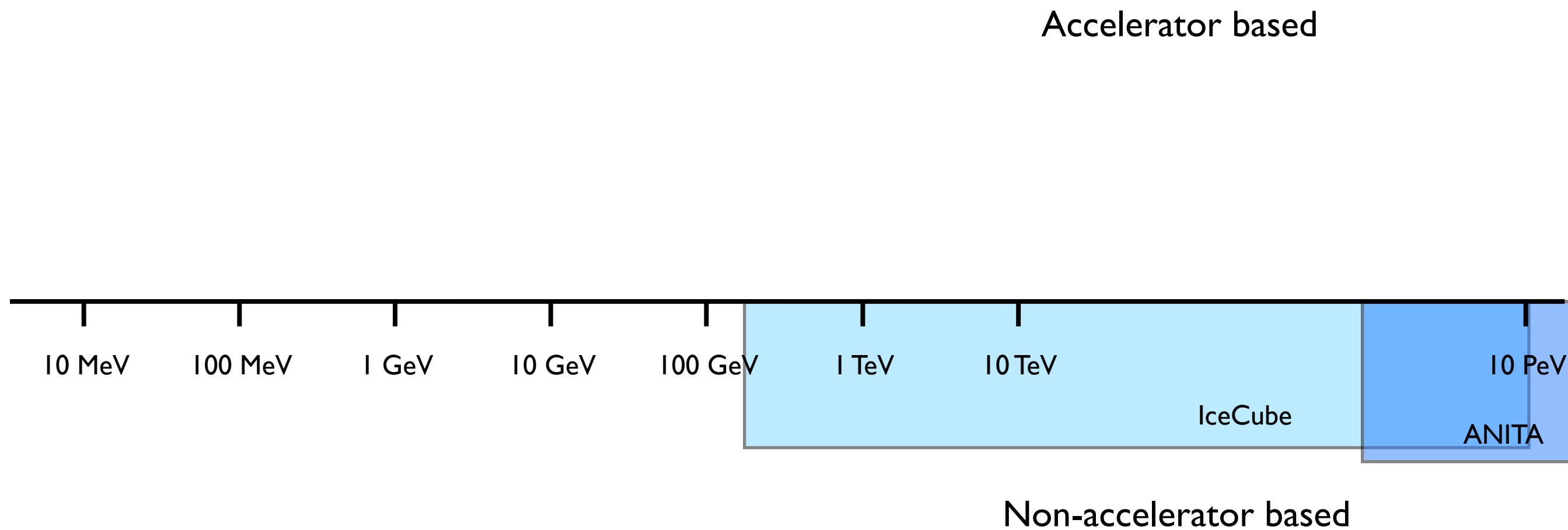
- Trigger level background to signal ratio is  $10^6:1$
- DeepCore uses IceCube as an active veto to reject down-going atmospheric muons and neutrinos
  - Atmospheric muon rejection of  $\sim 8 \times 10^3$  with neutrino retention of  $\sim 99\%$
  - Further rejection employed offline



# DeepCore Neutrinos

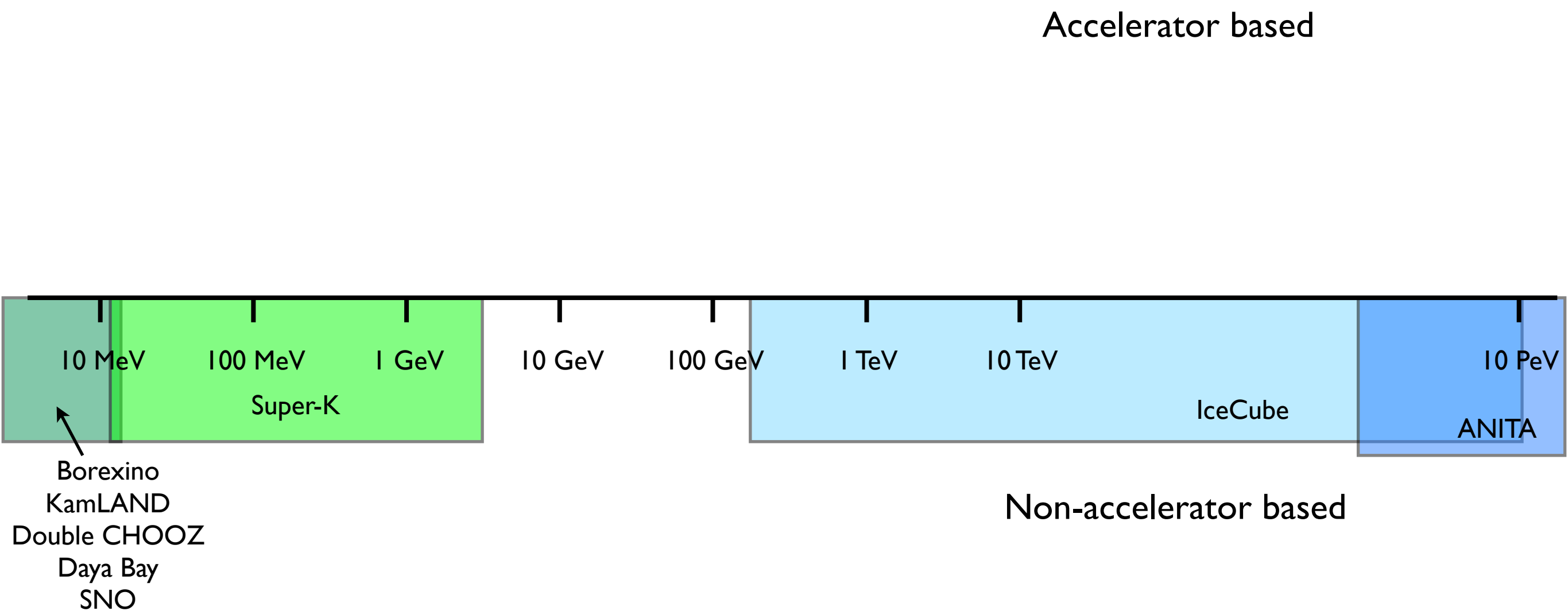
# Experimental Landscape

- IceCube
- DeepCore
- Beyond DeepCore



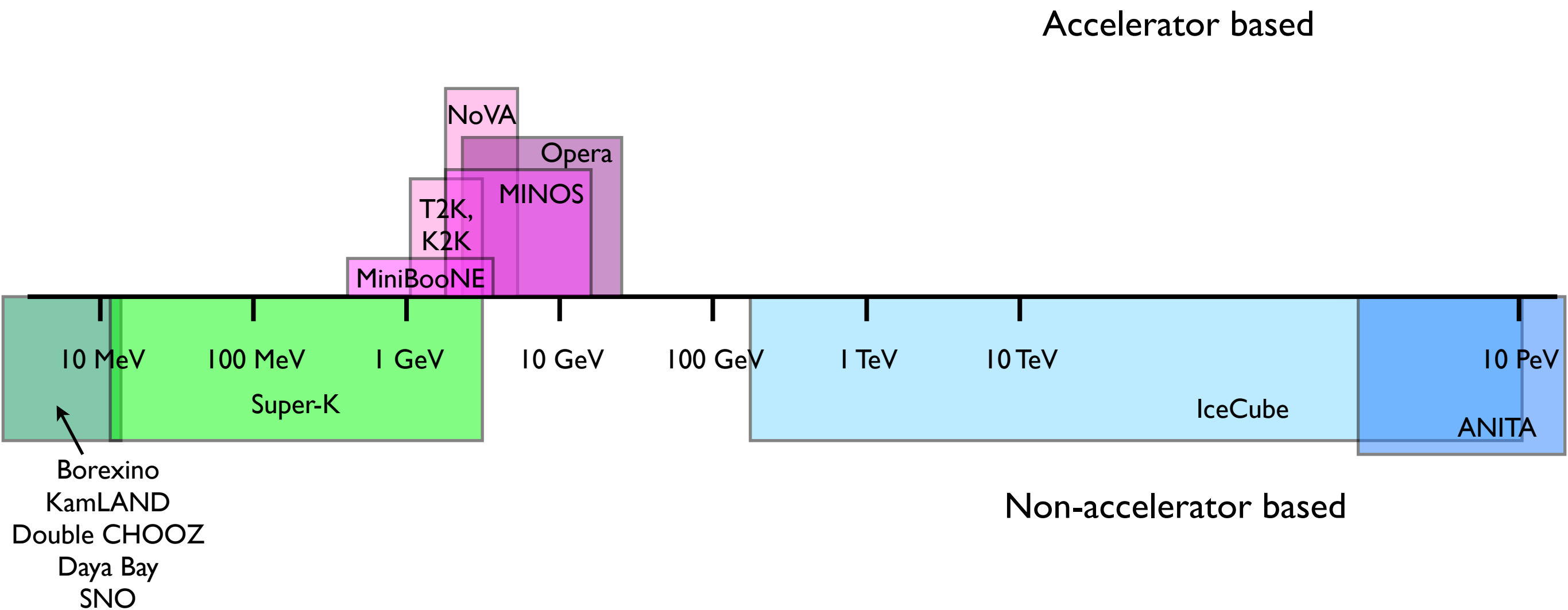
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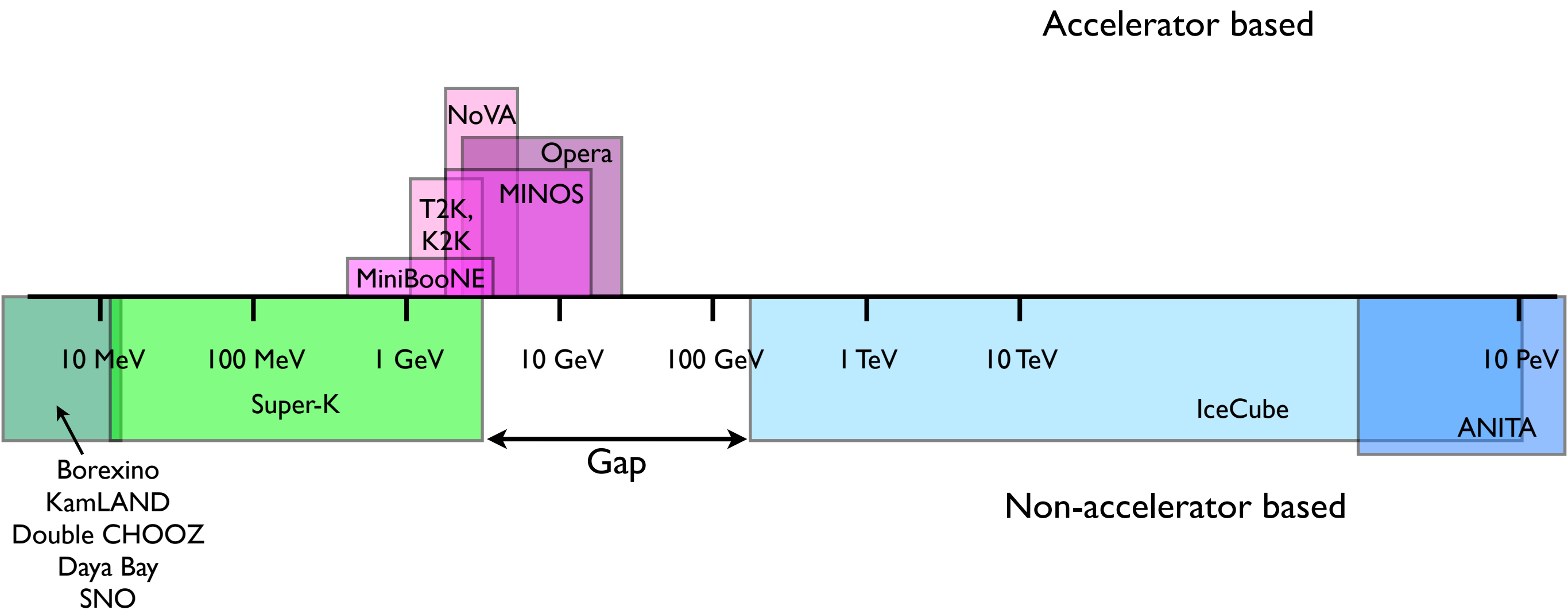
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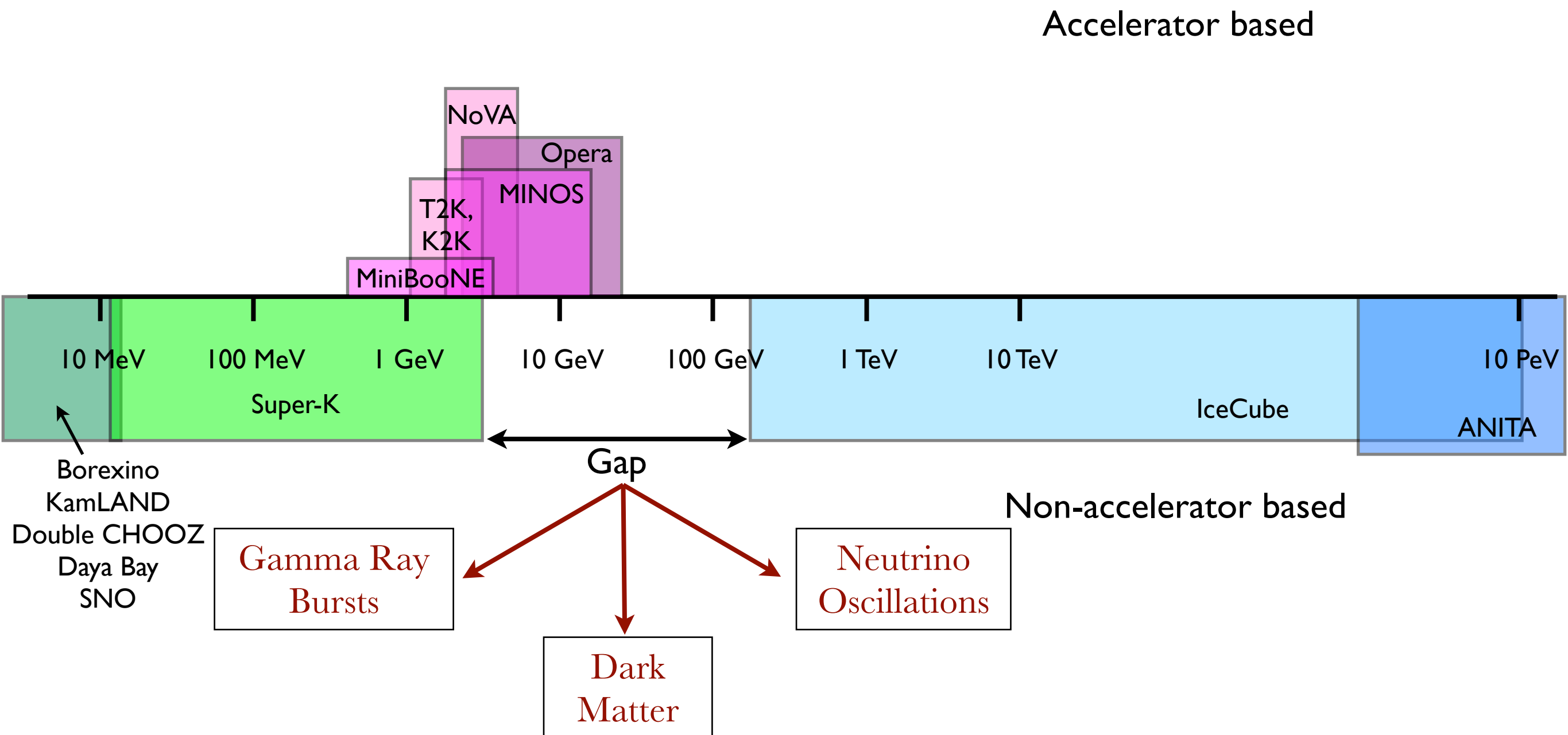
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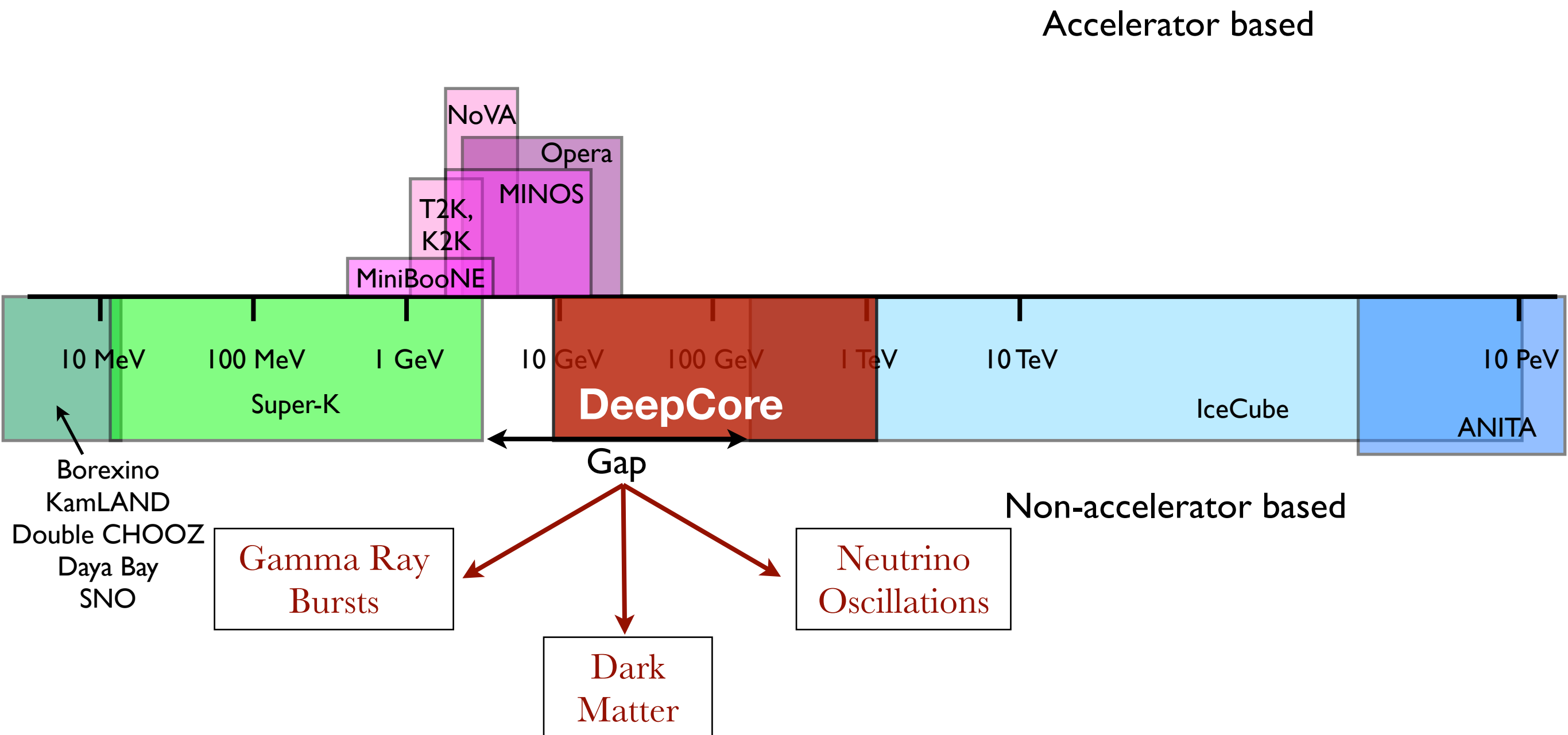
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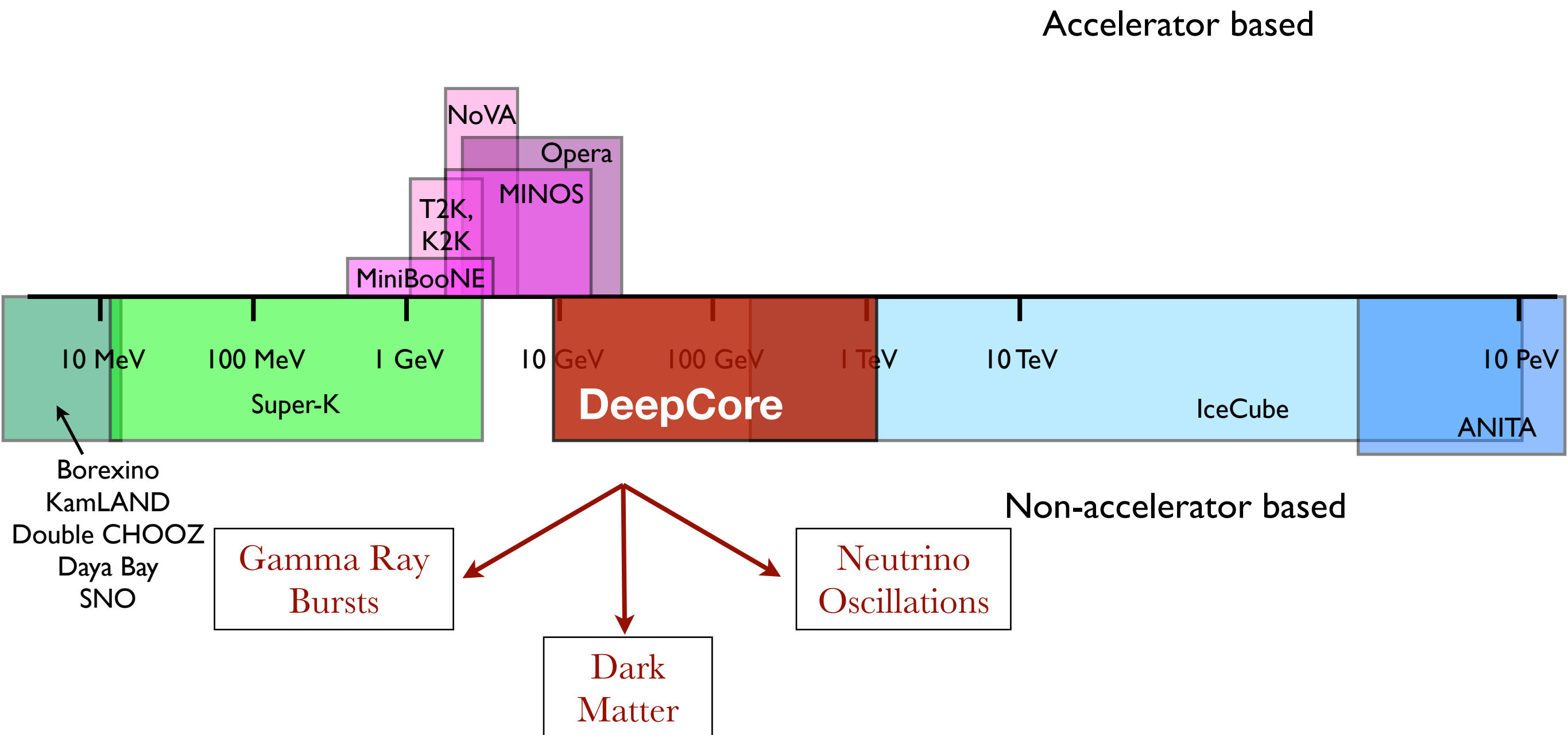
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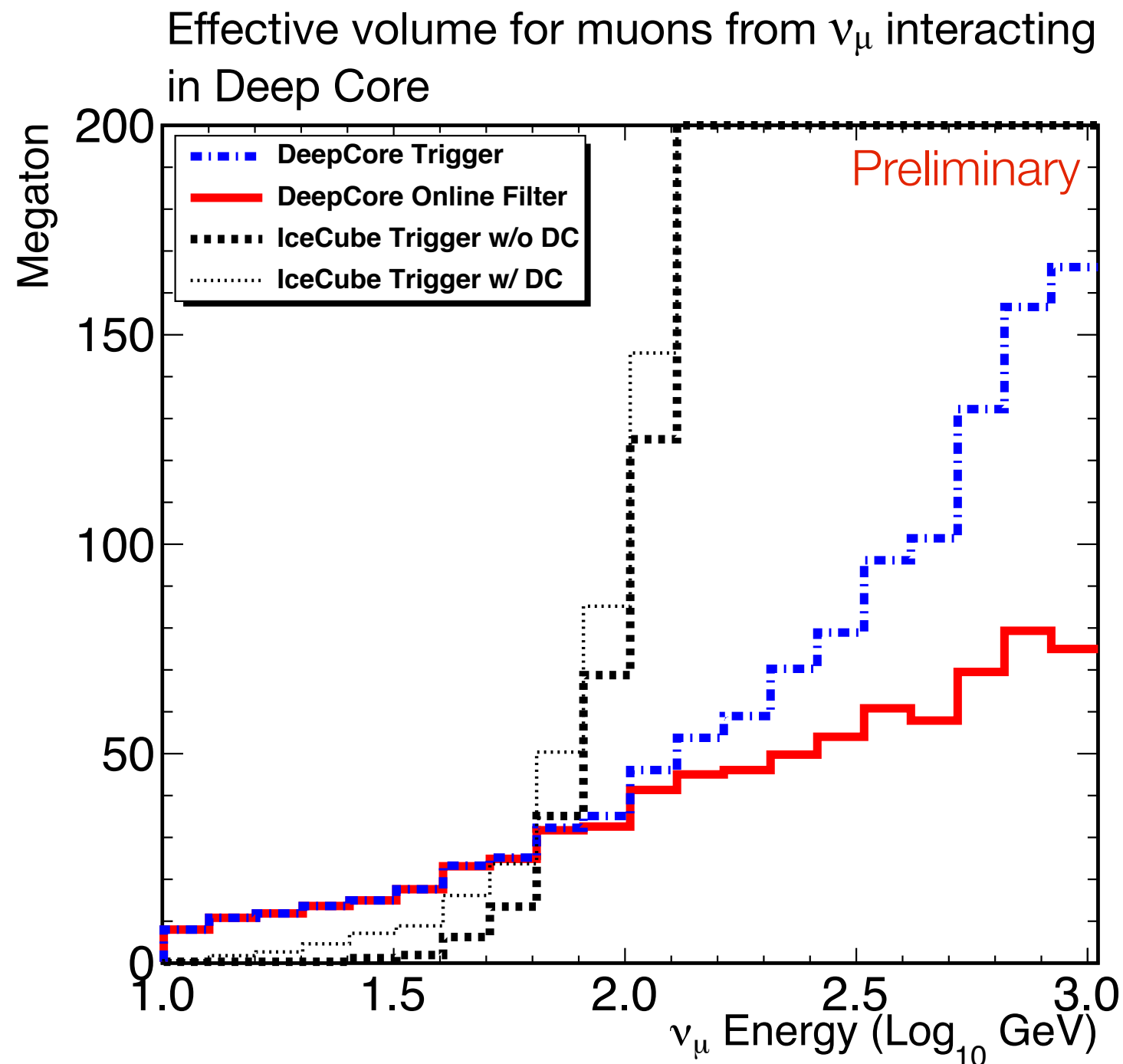


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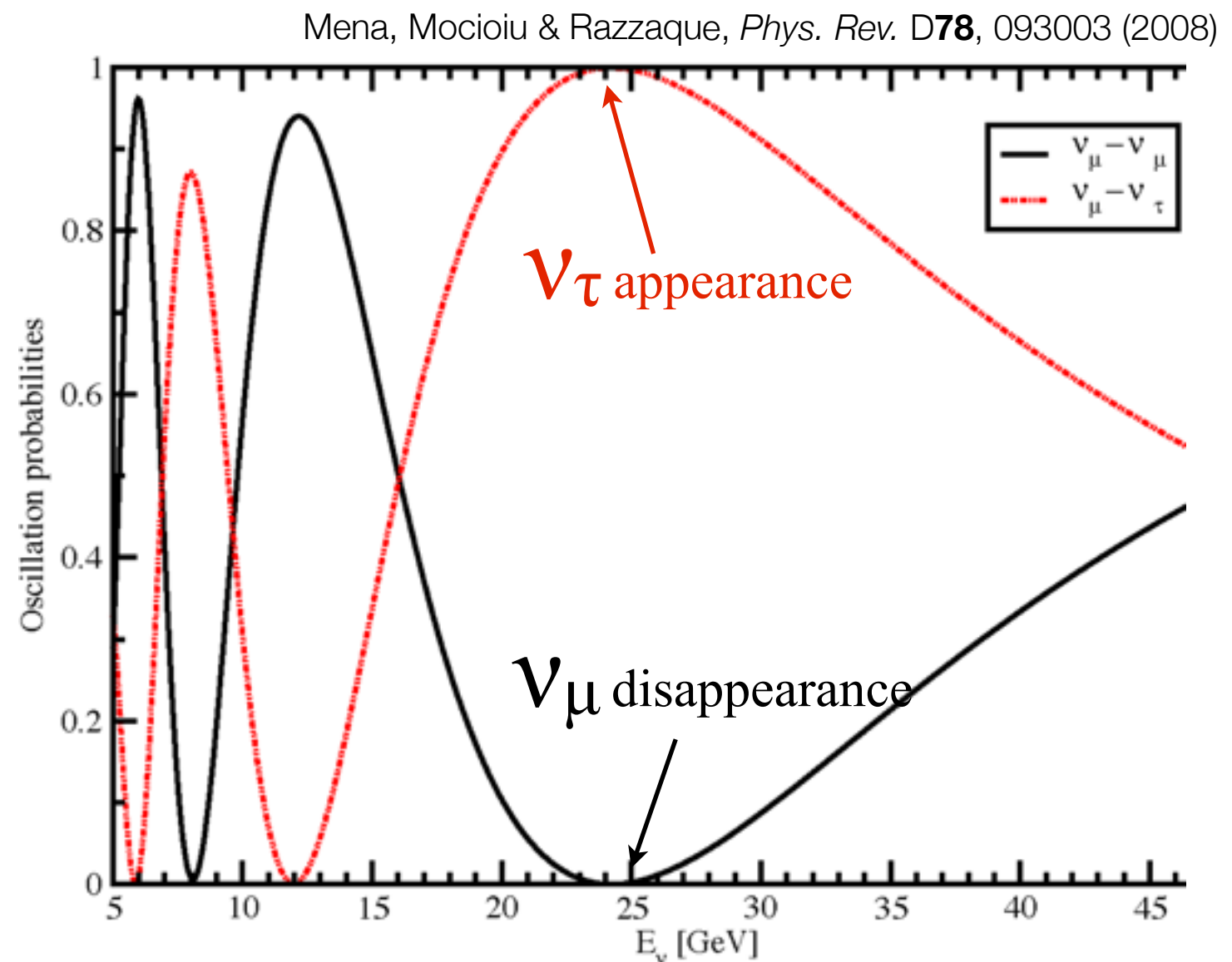
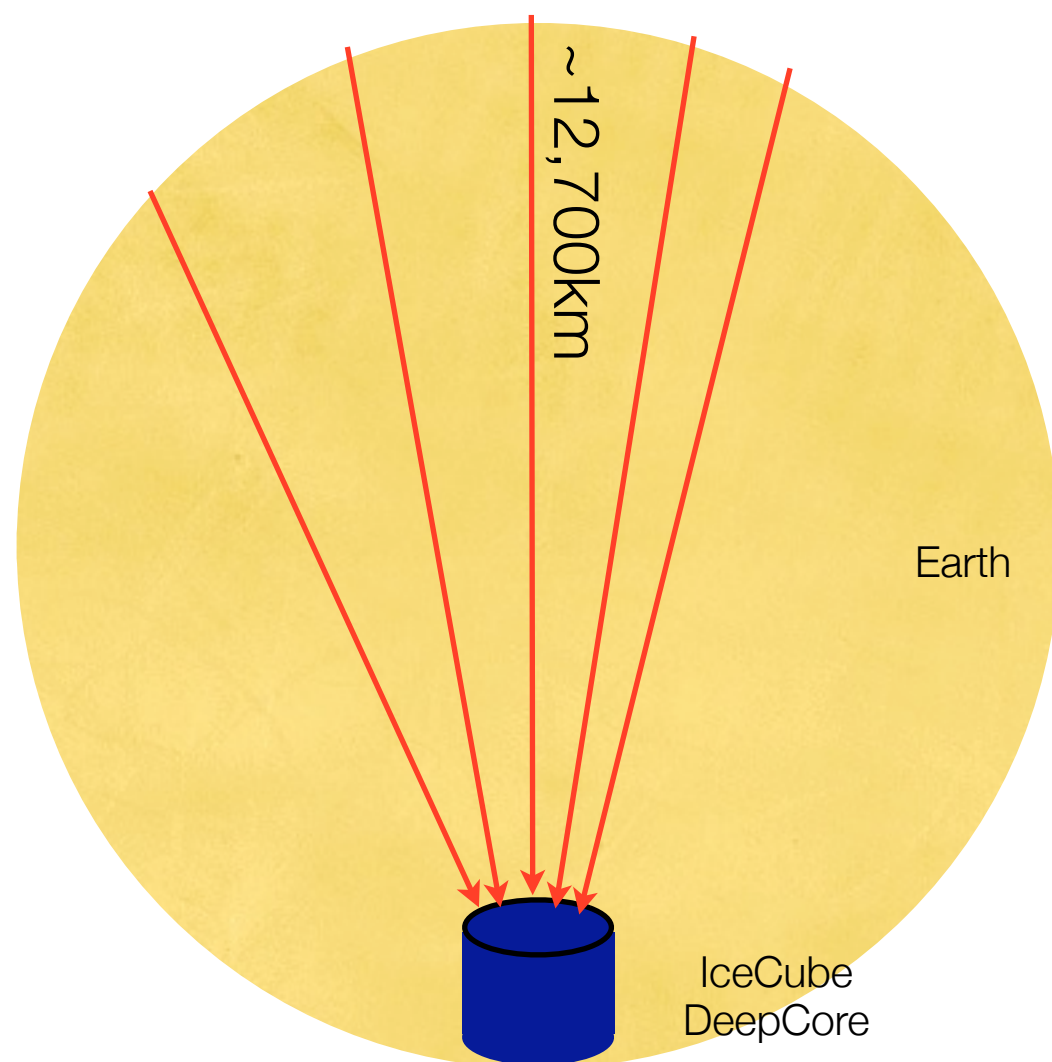
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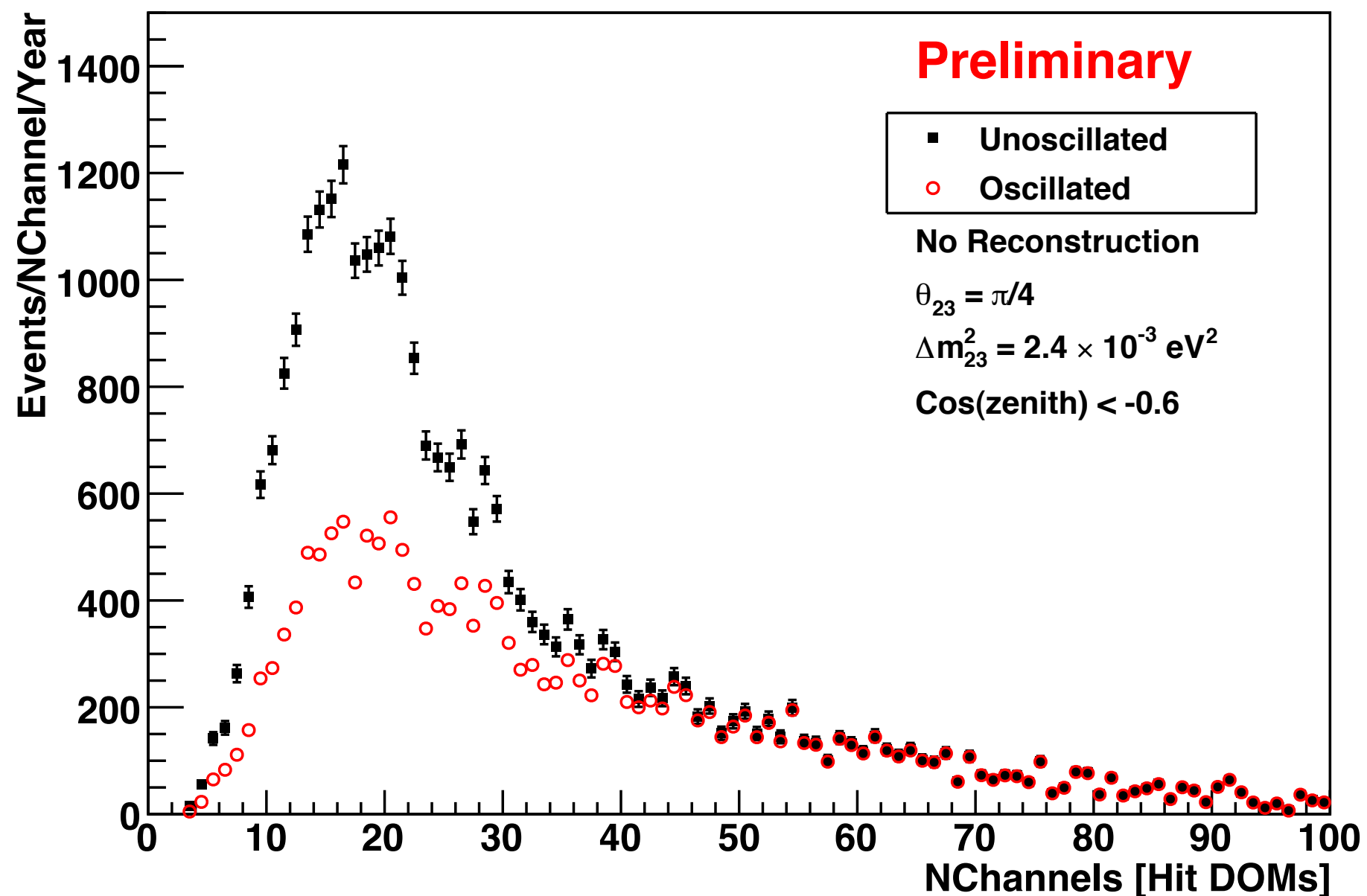
- DeepCore, being a multi-MTon detector, will collect tens of thousands of oscillated neutrinos at trigger level



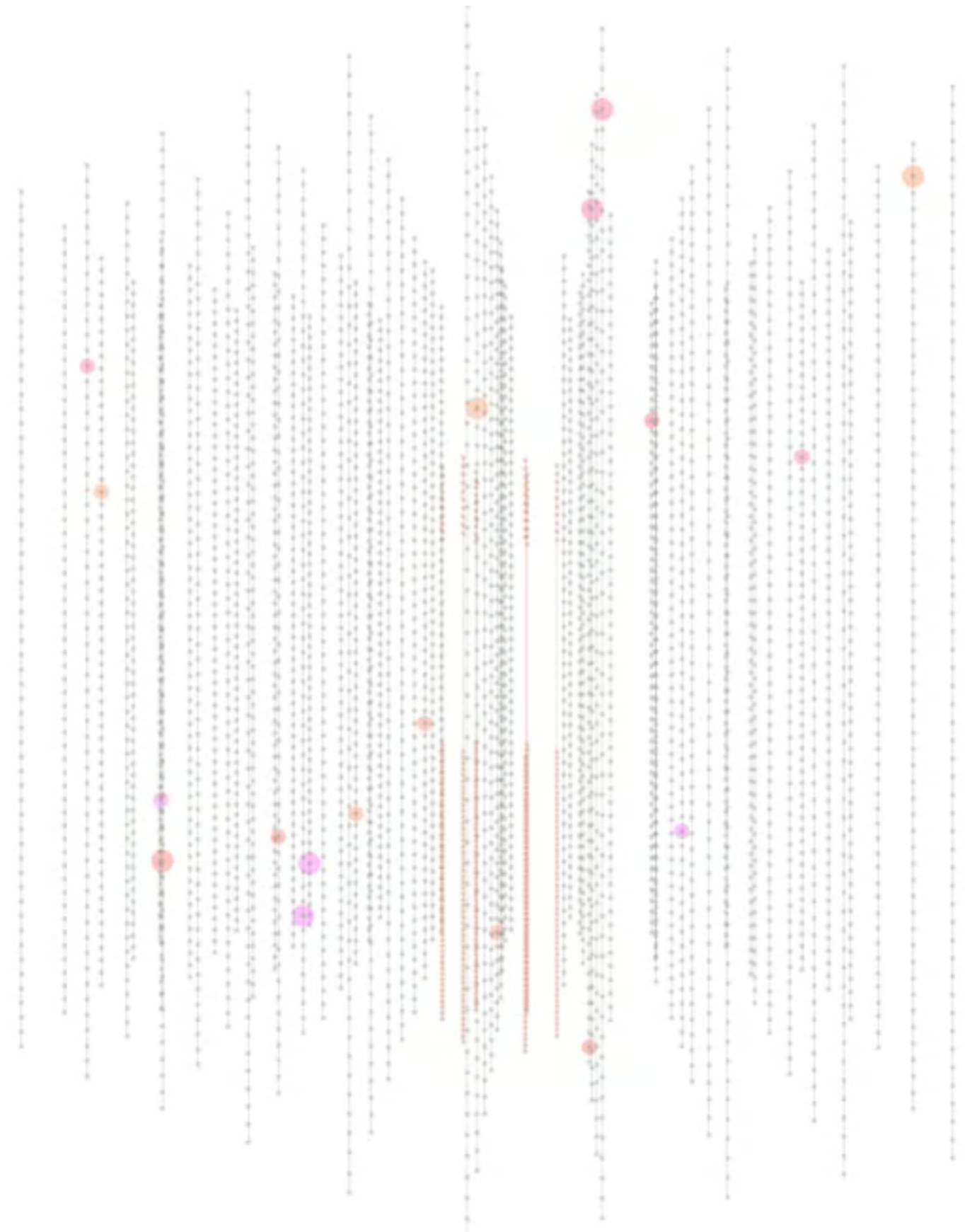
- Northern Hemisphere neutrinos oscillating over one earth radii produces  $\nu_\mu$  ( $\nu_\tau$ ) oscillation minimum(maximum) at  $\sim 25$  GeV
  - Higher energy region than accelerator based experiments
  - Beam never turns off



- 1 year data with 79 strings
- Monte Carlo signal only

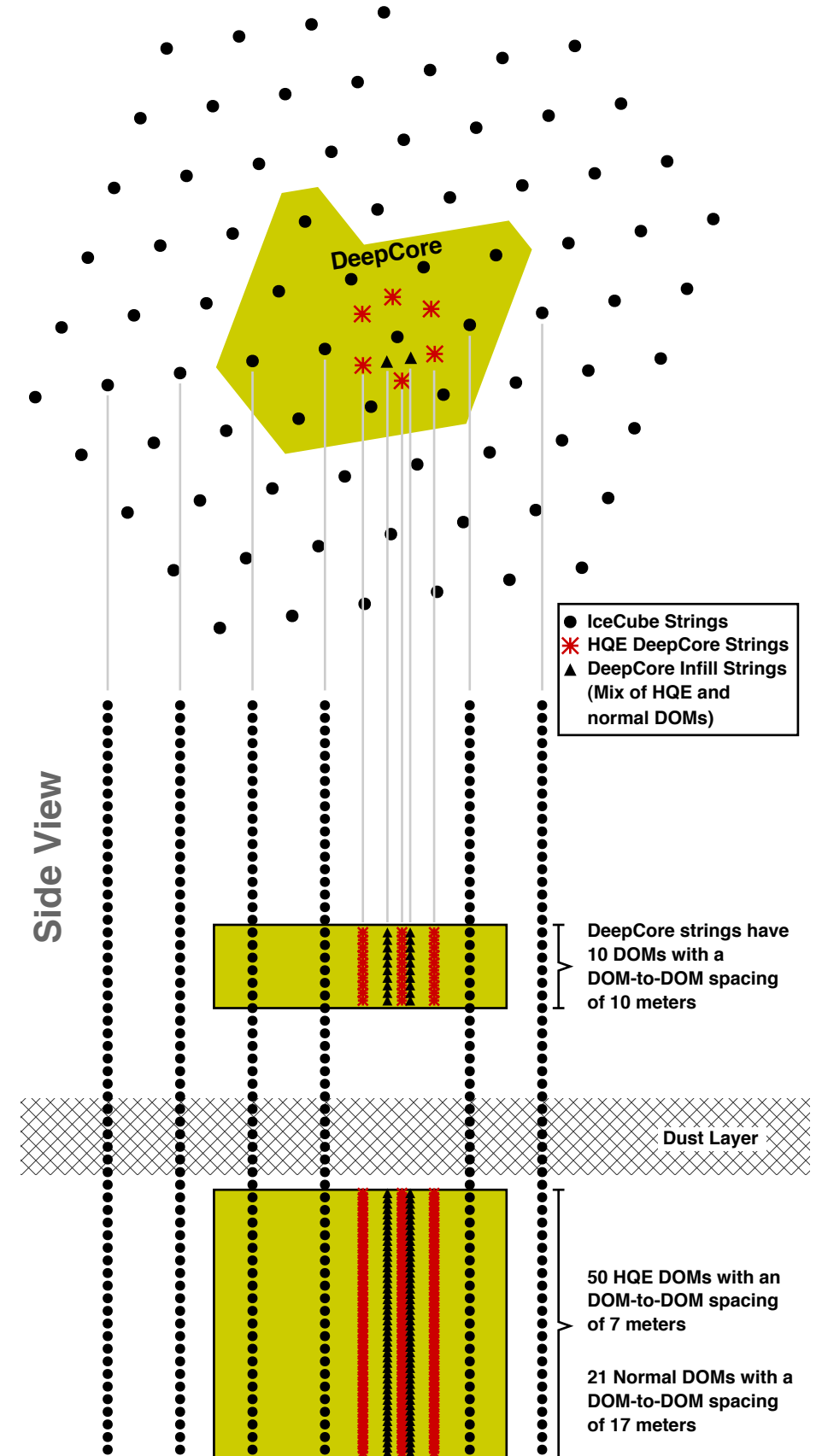


- 8 hours of real data
- Specific DeepCore vetoes and a Boosted Decision Tree
- Up-Going muon neutrino candidate
  - ~15 GeV from track length



- Neutral Current, Charged Current  $\nu_e$ , low energy  $\nu_\mu$  CC and CC  $\nu_\tau$  events produce cascade-like signatures
- Look for statistical excess in up-going cascade events
- DeepCore has been infilled with 2 additional strings
  - Increases  $\nu_\tau$  event rate by  $> 15\%$

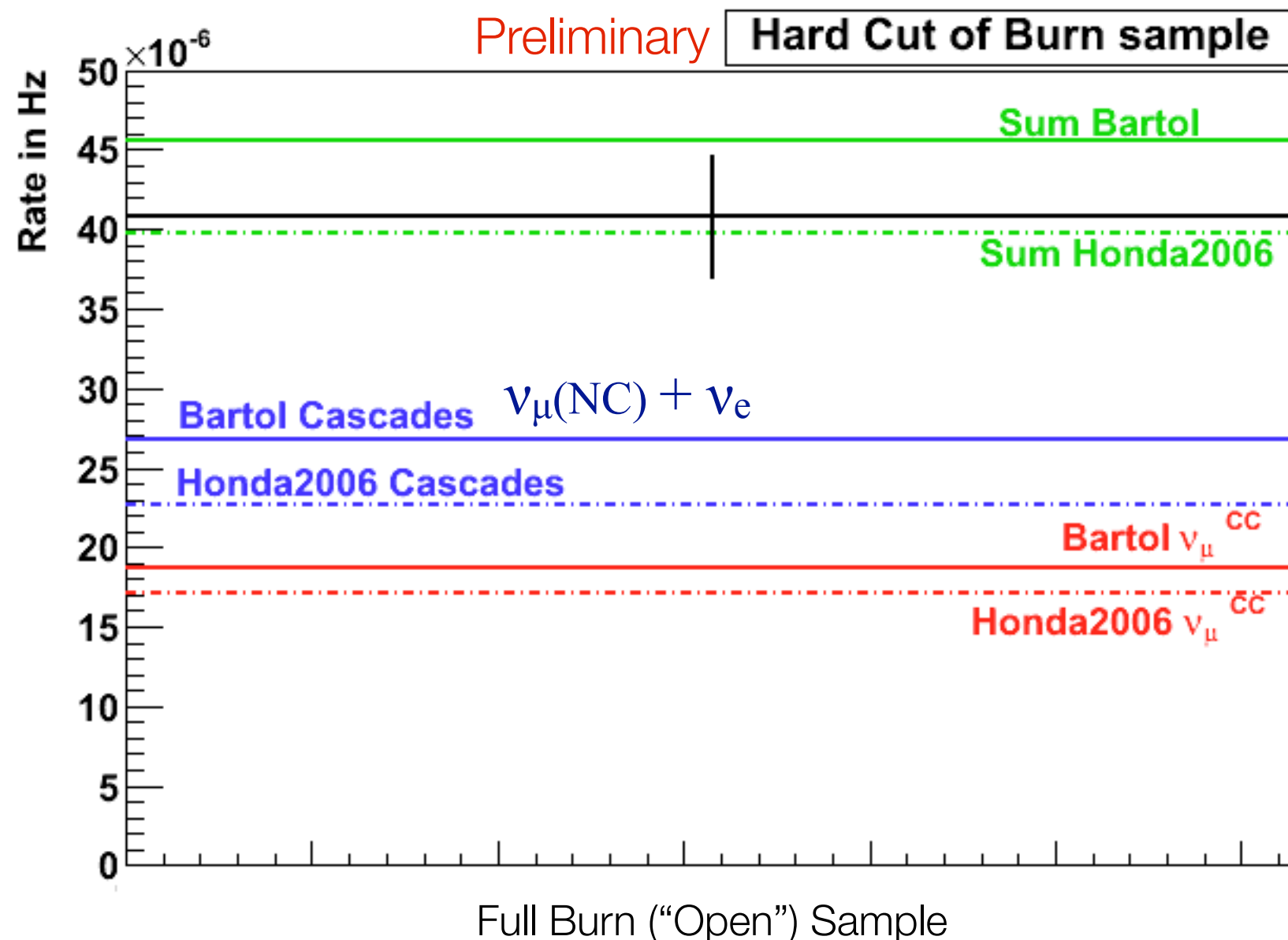
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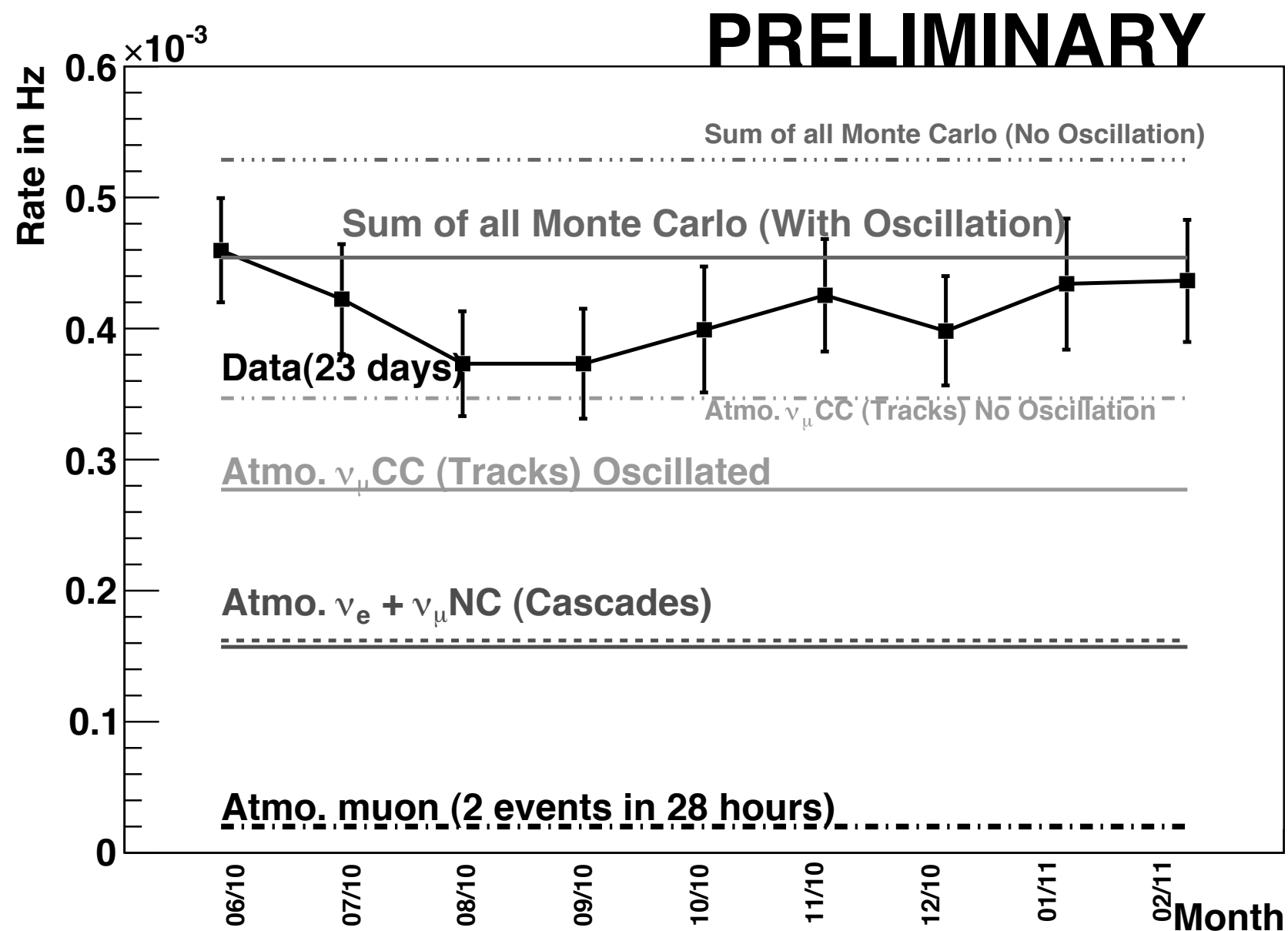
- DeepCore cascade candidate event



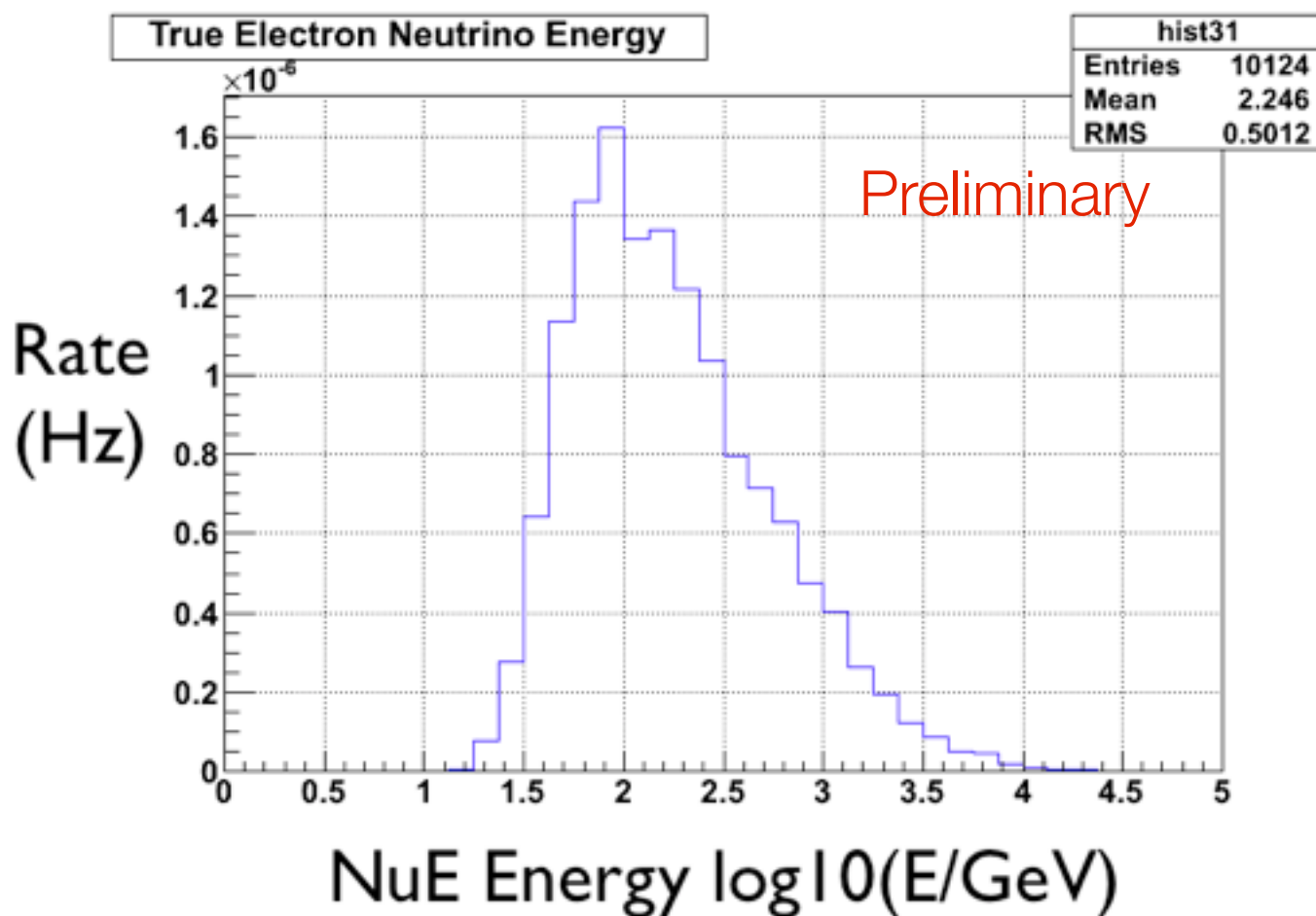
- First observation of neutrino induced cascades in IceCube
- 30.64 Days of IC79 livetime (10% open sample)
  - Tight cuts removed all cosmic ray muons from available 28 hours of CORSIKA
  - Inclusive rate over all



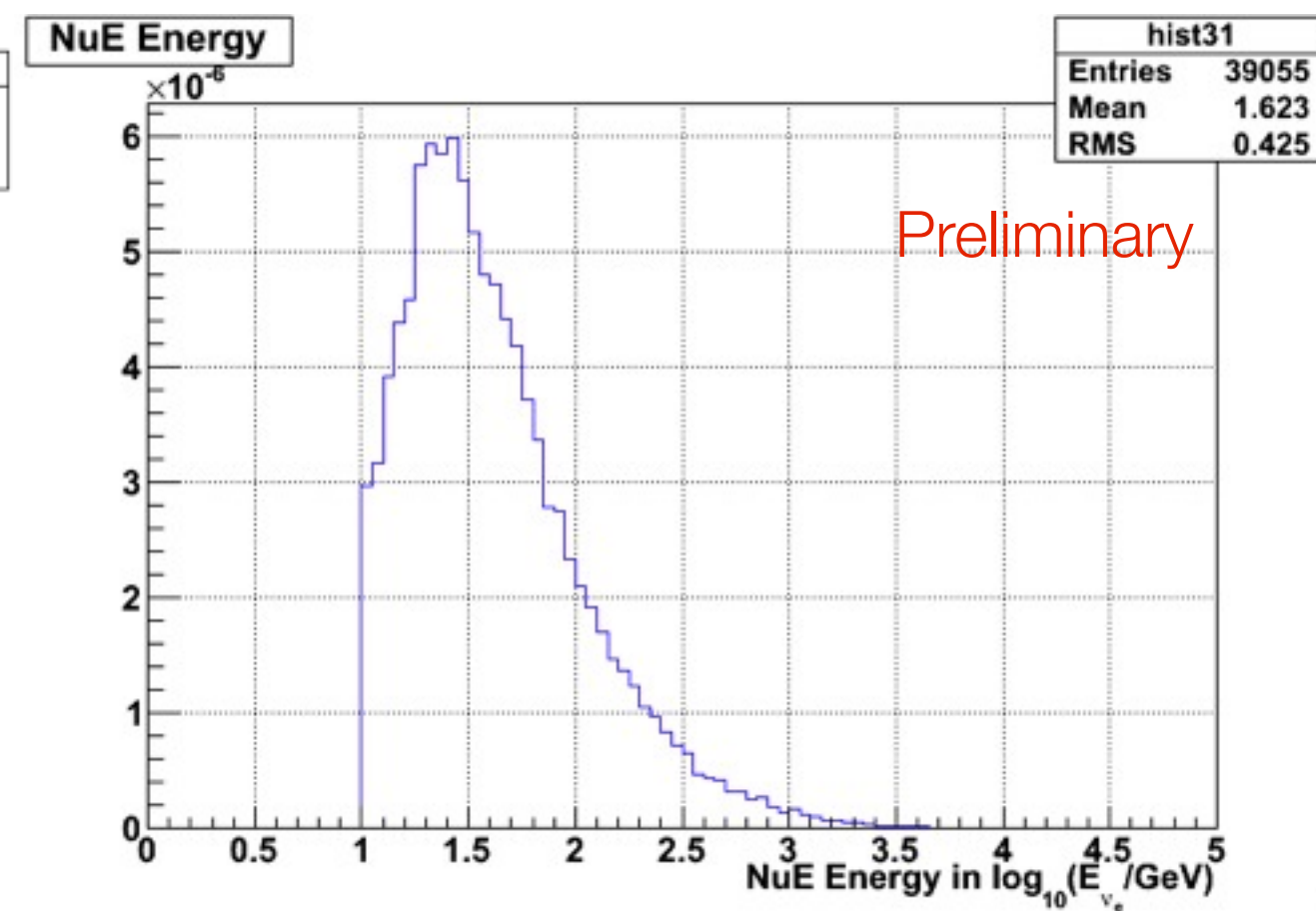
- Loose Cuts = More signal + non-zero atmospheric muon background
- Lower Energy Neutrinos



## Hard Cut



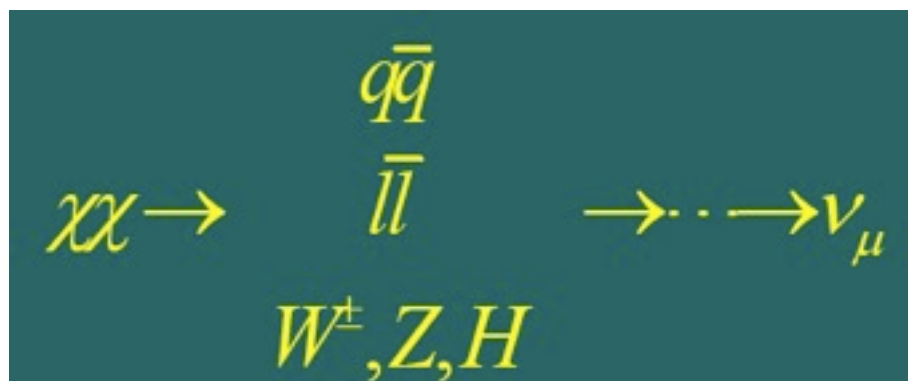
## Loose Cut



- 1 year of physics quality data with 80% deployment
- 100% deployed in Dec. 2010
- Opens up energy region for neutrino oscillation studies
- First observation of neutrino induced cascades

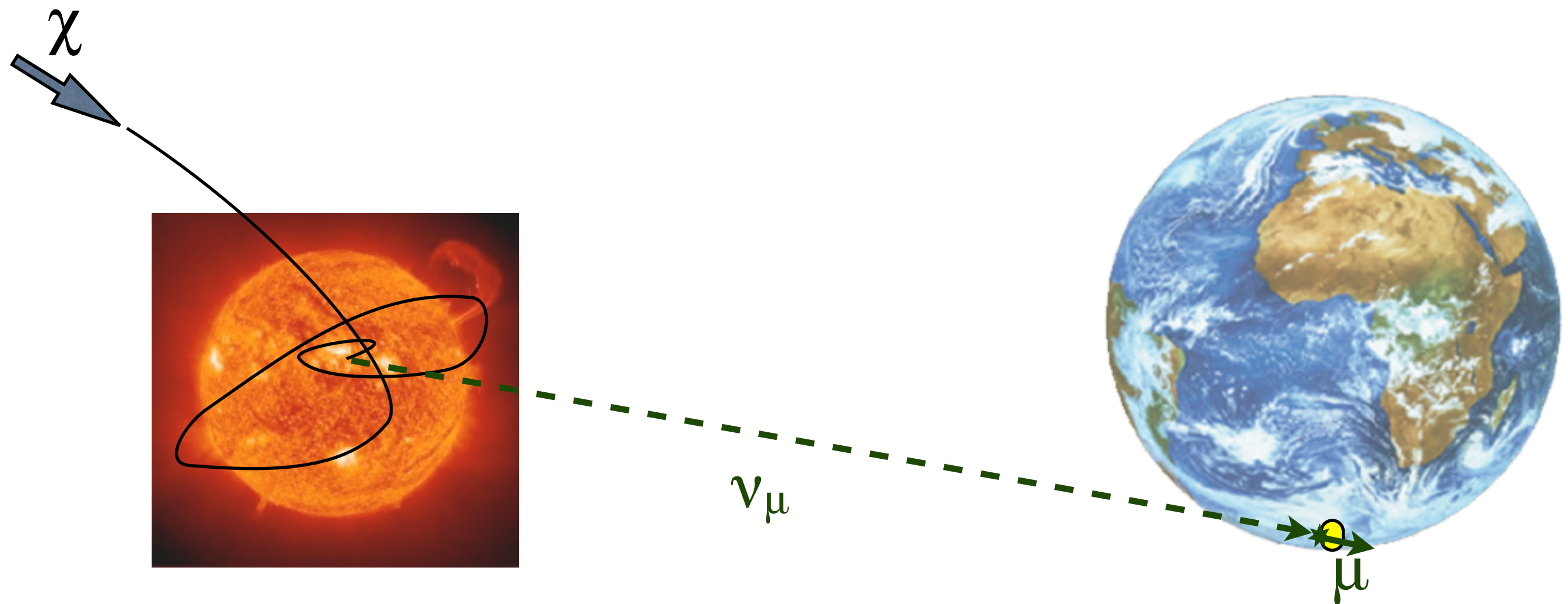
# DeepCore Dark Matter

- Galaxy clustering, Gravitational lensing, Bullet Cluster galaxies, etc... strongly suggest existence of Dark Matter
- Popular candidate for Dark Matter particle is **Weakly Interacting Massive Particle (WIMP)**
- IceCube-DeepCore searches for Dark Matter self-annihilation creating neutrinos
  - Point towards galactic objects where Dark Matter clumps ( Sun, Galactic Halo)

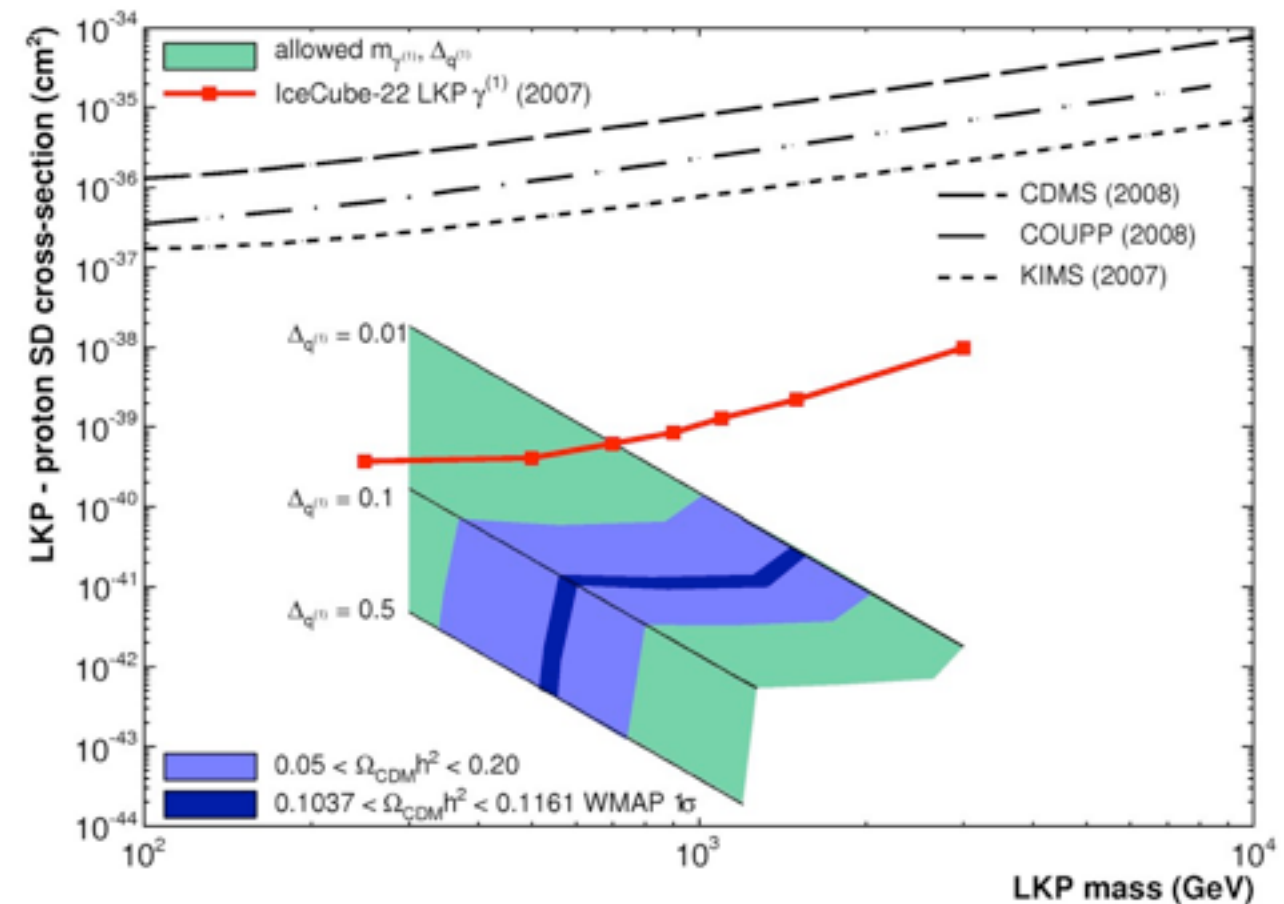
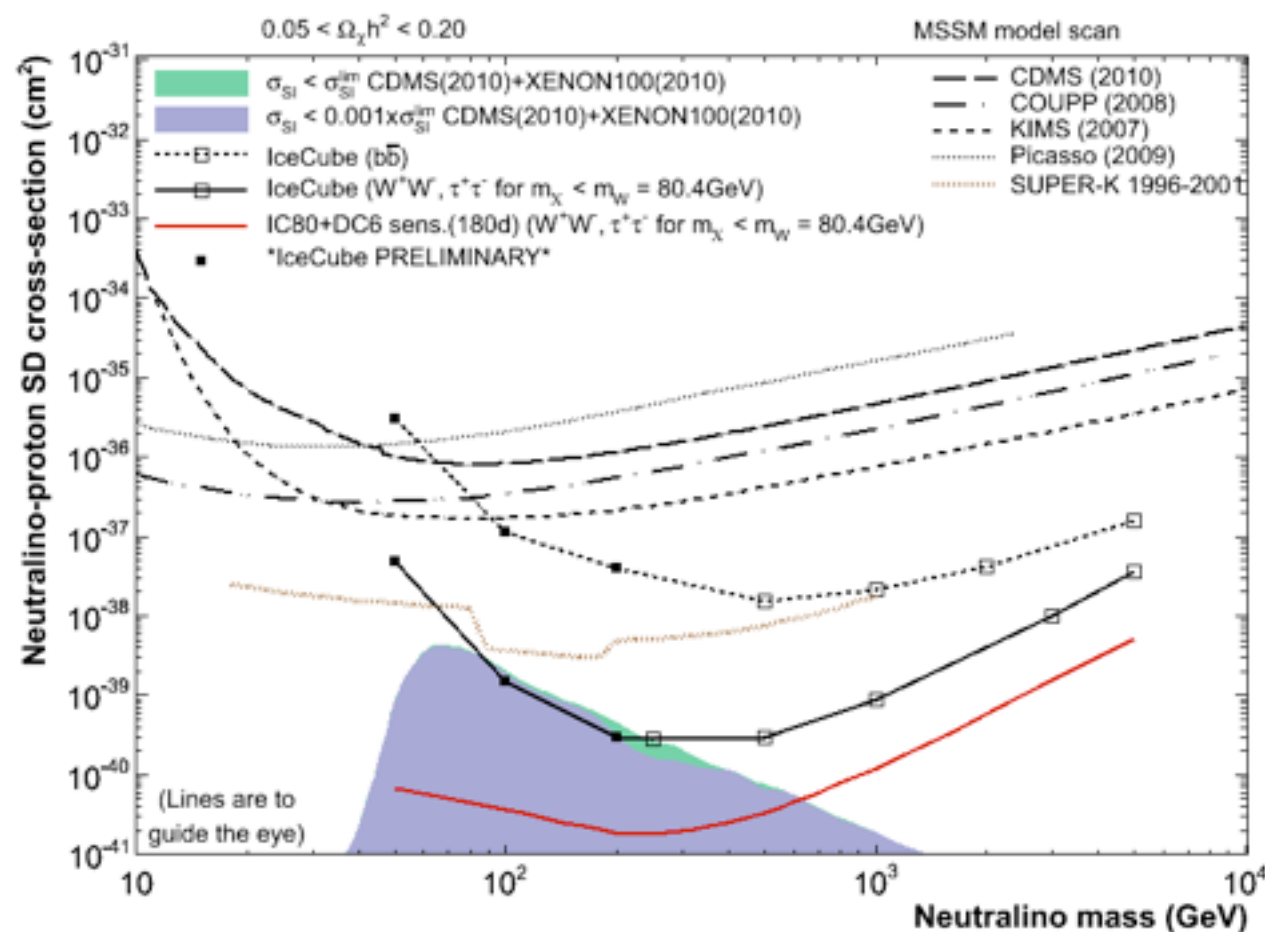


W. H. Press and D. N. Spergel. *Astrophys. J.* **296**, 679, (1985)  
 T. Gaisser, G. Steigman and S. Tilav. *Phys. Rev. D* **34**, 2206, (1986)  
 A. Gould. *Ap. J.* **328**, 919, (1988).

- Dark Matter gets trapped via Spin-Dependent (SD) collisions with nucleons
  - Spin-independent is well constrained by direct detection experiments
- Builds up, reaches equilibrium, starts annihilating



- Look for neutrino excess when the Sun is below the horizon
- Neutrino flux translates to cross-section
- DeepCore will provide order of magnitude+ improvement in spin dependent searches for Dark Matter
- Limits on MSSM model Dark Matter and Kaluza-Klein model



- Current IceCube-only analyses of spin dependent WIMP are probing phase space
- Inclusion of DeepCore will improve WIMP Dark Matter sensitivity by 2 orders of magnitude

# Beyond DeepCore



IceCube



DeepCore

# Beyond DeepCore



IceCube



DeepCore

# Beyond DeepCore



IceCube



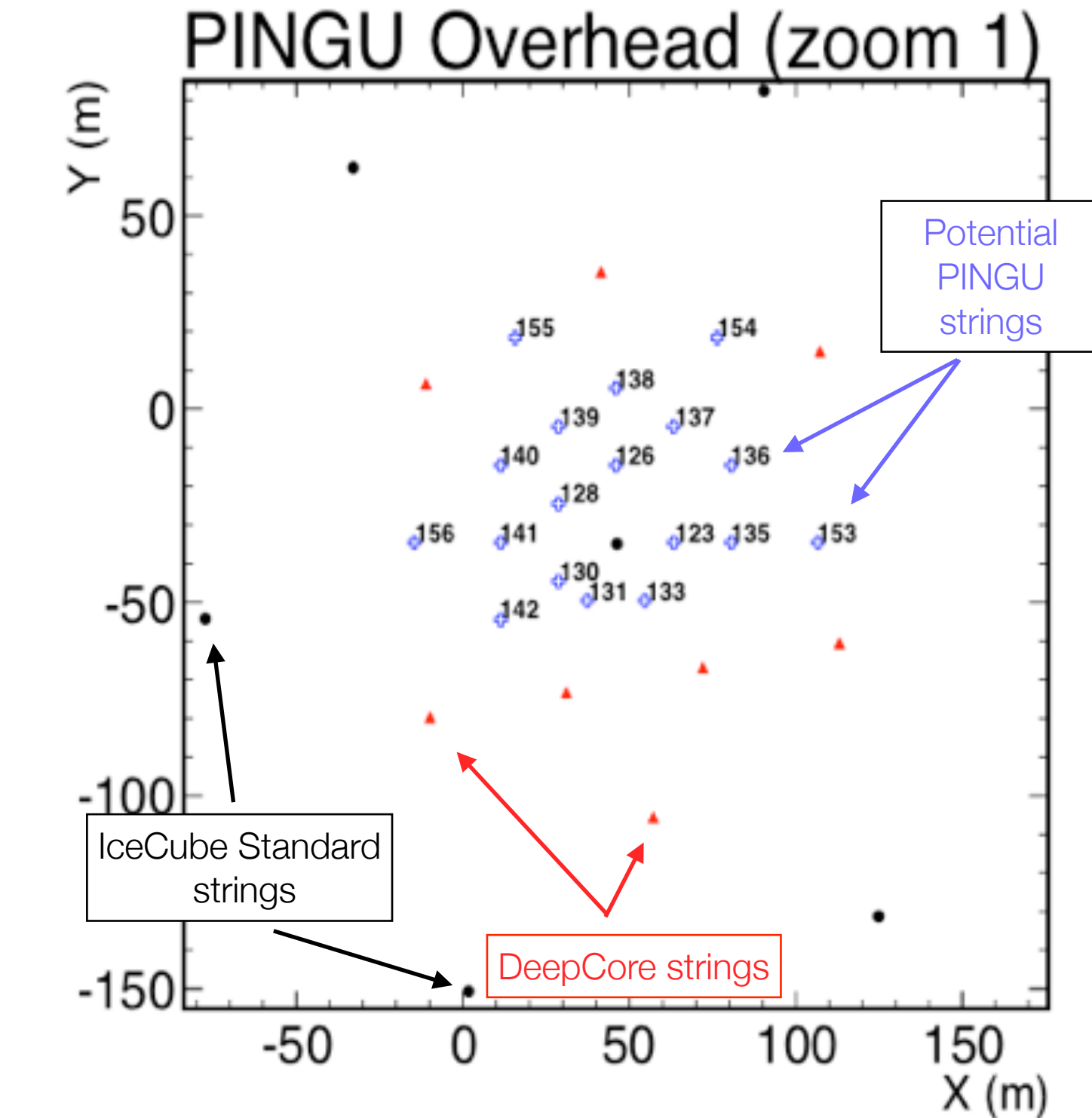
DeepCore



PINGU

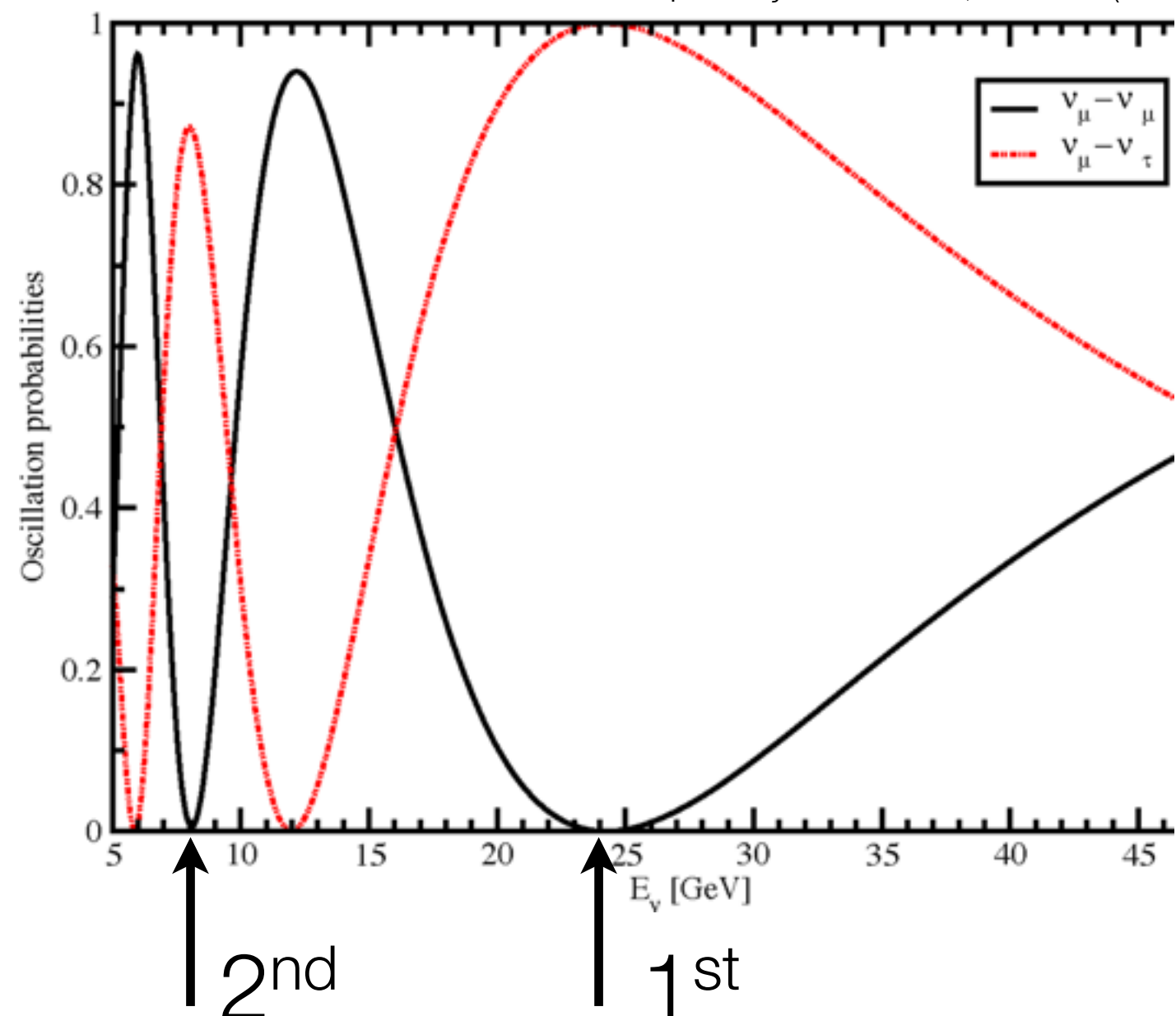
- Using existing and familiar technology ( hot water drill, HQE PMT DOMs) infill DeepCore with 18-20 more Strings
- Drive neutrino energy reach down to few GeV while maintaining multi-megaton scale size
- Improves Dark Matter sensitivity and Neutrino Oscillation analyses
- Phased IceCube Next Generation Upgrade (PINGU)

- Add 18-20 strings into DeepCore volume
- One of many possible geometries
- R & D for future water/ice cerenkov detectors



- With an infill that achieves  $\sim \text{GeV}$  resolution, the 2nd oscillation minimum becomes accessible
- Improve Cascade reconstruction
  - Tau appearance

Mena, Mocioiu & Razzaque, *Phys. Rev. D***78**, 093003 (2008)



- Monte Carlo is now available using GENIE neutrino generator for GeV neutrinos
- Soliciting interest in those who want to see a megaton water cerenkov detector at GeV neutrino energies
- <http://www.mailman.srv.ualberta.ca/mailman/listinfo/beyonddc>

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**“Anything worth  
doing is worth  
overdoing”**

# Really Beyond DeepCore



IceCube



DeepCore



PINGU

# Really Beyond DeepCore



IceCube



DeepCore



PINGU

# Really Beyond DeepCore



IceCube



DeepCore



PINGU



- South Pole Infrastructure
  - No excavation
  - Deployment is now a precision process
- Unchanging, low-background medium
- Use IceCube and DeepCore (maybe PINGU) as active veto
- Move from GeV to MeV
  - Cerenkov Ring Imaging
  - Single PMT Module is no longer feasible

- Based on a KM3NeT proposed design

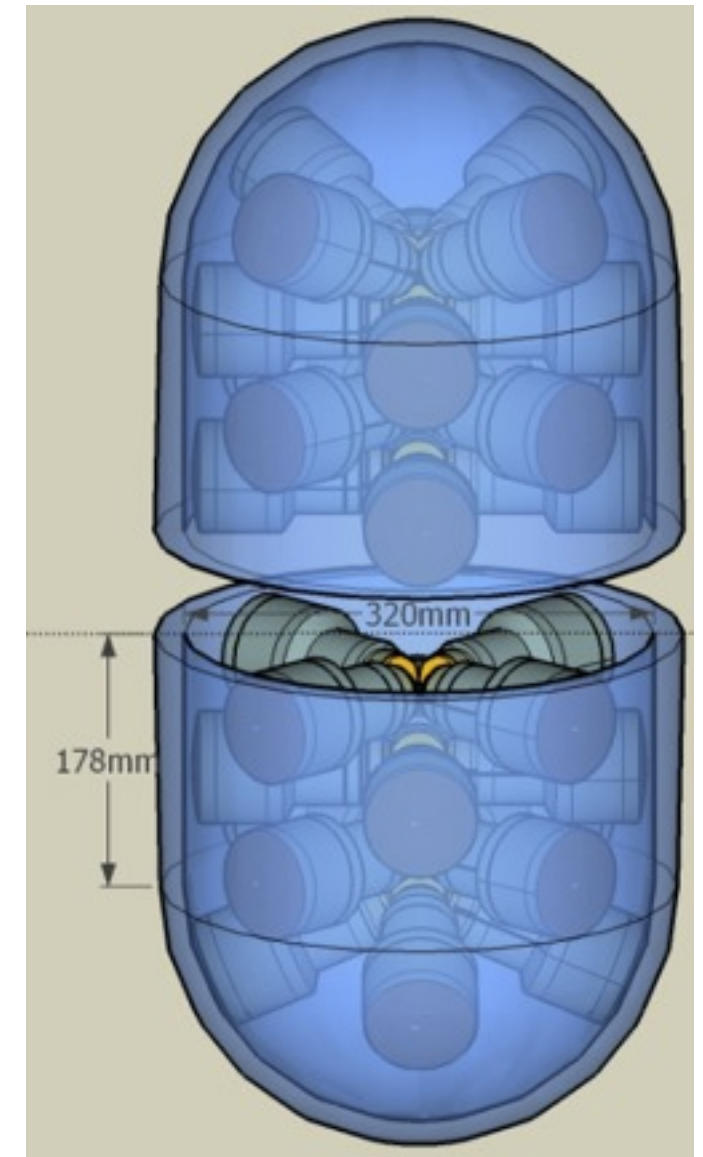
P. Kooijman, *NIM A567* (2006), S. Kuch *NIM A567* (2006), KM3NeT TDR

- One meter glass cylinder containing 30 3" PMTs and associated electronics

- Comparable width to IceCube DOM
- Effective photocathode area of 265 sq. in. – 3.4x that of standard 10" IceCube PMT, but granular

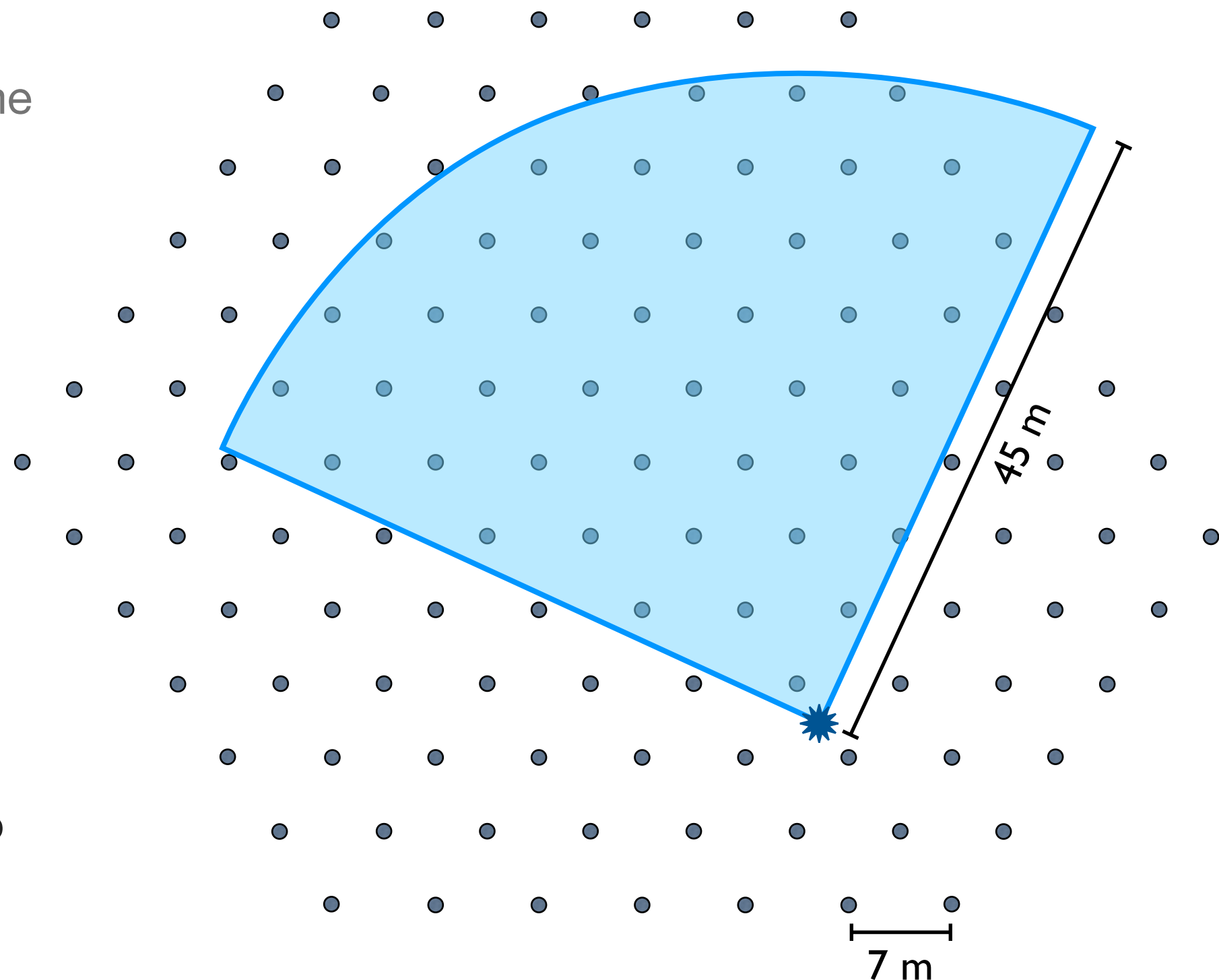
- Would allow spatial imaging of Cherenkov ring

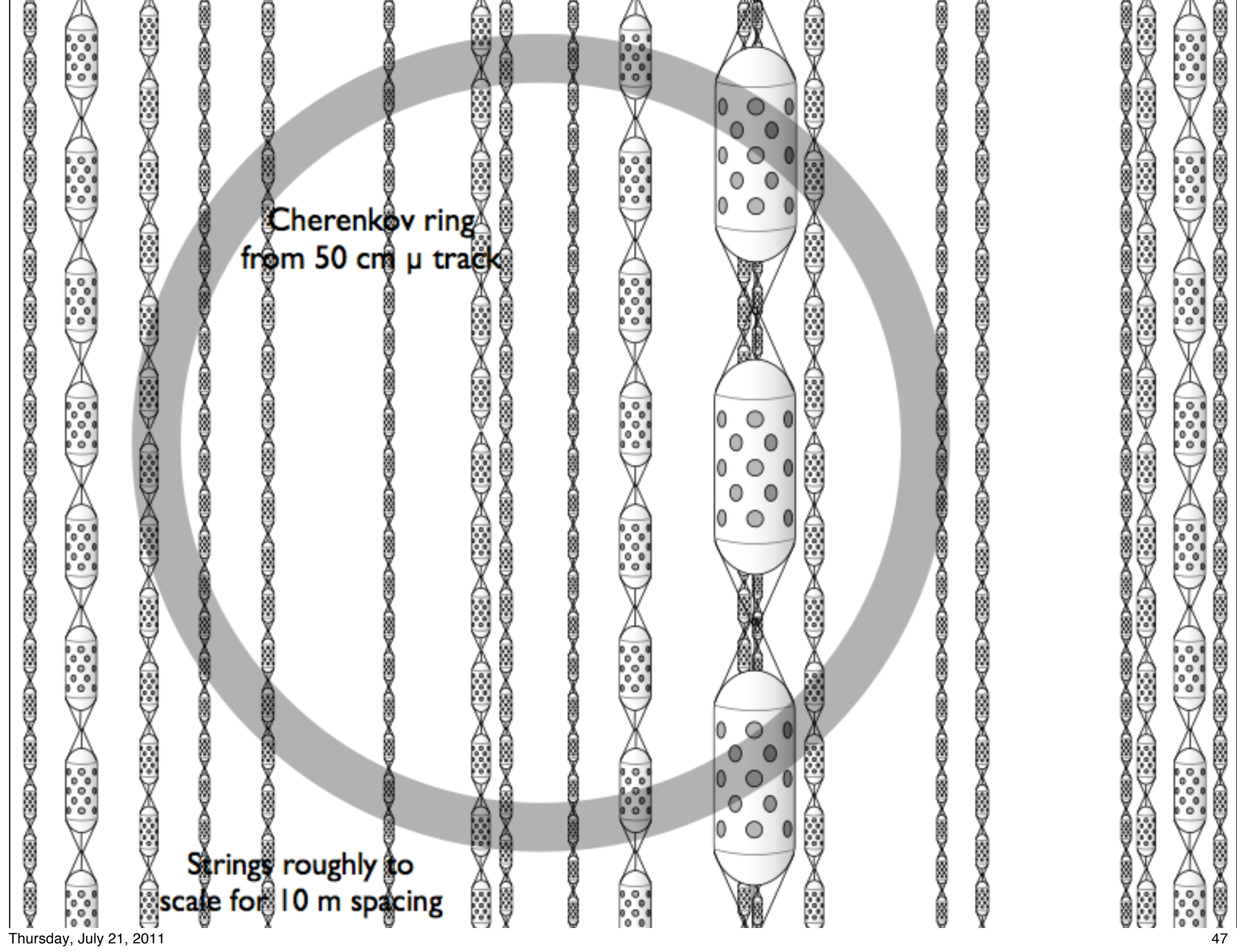
- KM3NeT has moved toward a 17" sphere instead – close PMT spacing not their goal – but the cylinder design was developed to the engineering stage



Courtesy P. Kooijman

- 120 strings of 125 composite DOMs each
  - Instrumented volume of 250 m height, ~40 m radius
- 1 MegaTon fiducial volume, at depths of 2200-2450 m
  - Optical scattering length  $\gtrsim 40$  m, absorption  $\gtrsim 140$  m
- Inside IceCube and DeepCore for muon veto





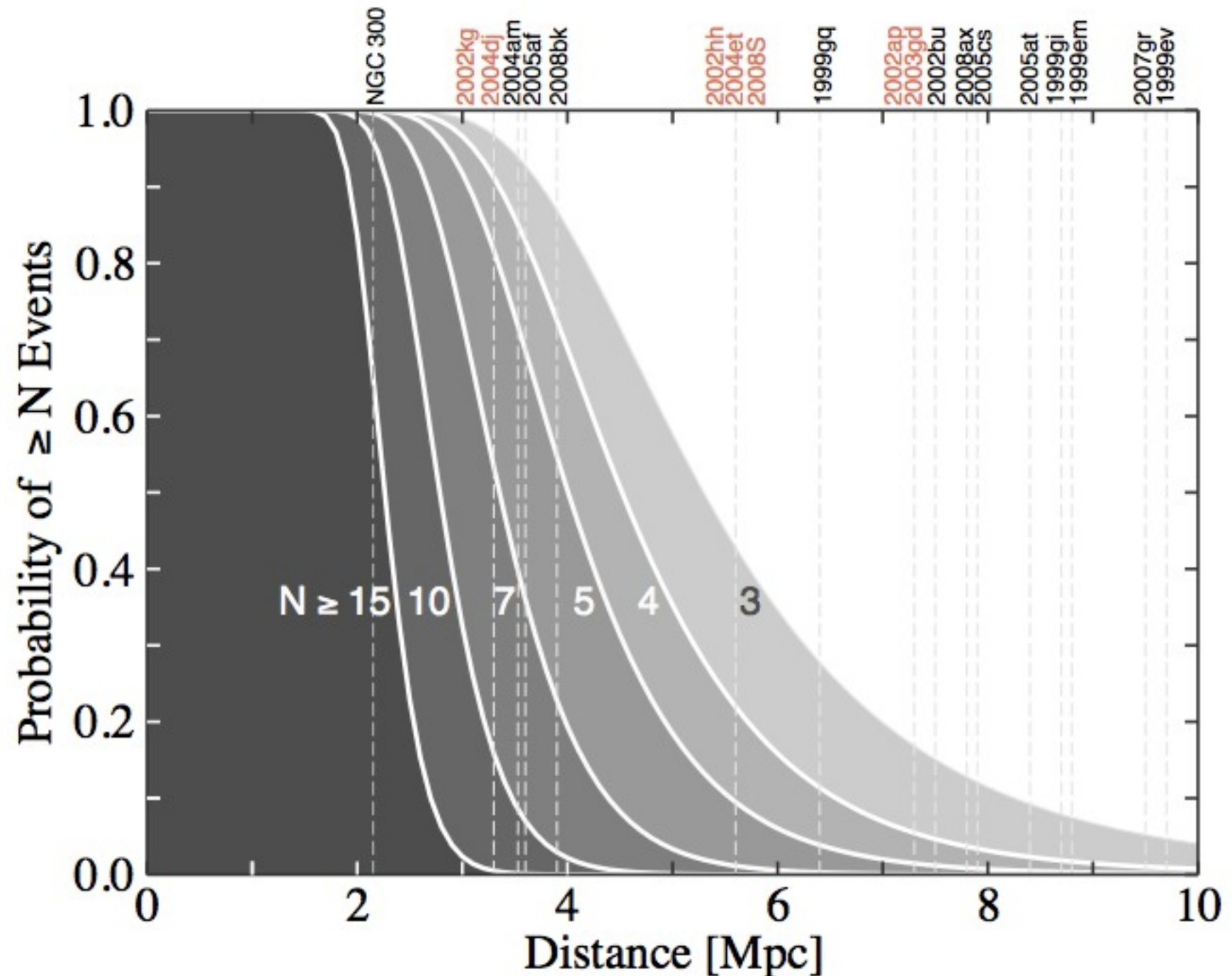
The diagram illustrates a detector setup with vertical strings. A central vertical string is highlighted with a larger, shaded cylindrical volume representing a 50 cm muon track. A large, faint gray circle is centered on this track, representing the Cherenkov ring. The other strings are shown as thinner, unshaded cylinders. The strings are connected by a network of lines, suggesting a readout system. The text 'Cherenkov ring from 50 cm  $\mu$  track' is placed within the gray circle. The text 'Strings roughly to scale for 10 m spacing' is located at the bottom left.

Cherenkov ring  
from 50 cm  $\mu$  track

Strings roughly to  
scale for 10 m spacing

- Extend core-collapse SN search beyond Milky Way
- 5 megaton detector with sensitivity down to 15 MeV

Kistler et. al.  
arXiv:0810.1959





- Detector for Neutrino Factory, Beta beam or Super Beam
  - Mass Hierarchy, Small  $\theta_{13}$  ( $10^{-4}$ - $10^{-3}$ ), Lepton CP Violation
  - Option for PINGU as well depending on beam characteristics

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- Proton Decay

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  - Mass Hierarchy, Small  $\theta_{13}$  ( $10^{-4}$ - $10^{-3}$ ), Lepton CP Violation
  - Option for PINGU as well depending on beam characteristics
- Proton Decay
- Extra-galactic Supernova neutrino

- IceCube astrophysical neutrino searches are alive and well
- DeepCore extends IceCube sensitivity to neutrinos down to  $\sim 10$  GeV
- Phased extensions aim to provide lower energy reaches at megaton sizes in a water(ice) cerenkov detector
  - PINGU -  $\sim 1$  GeV at trigger
  - Phase 2 - Down to  $\sim O(10)$  MeV

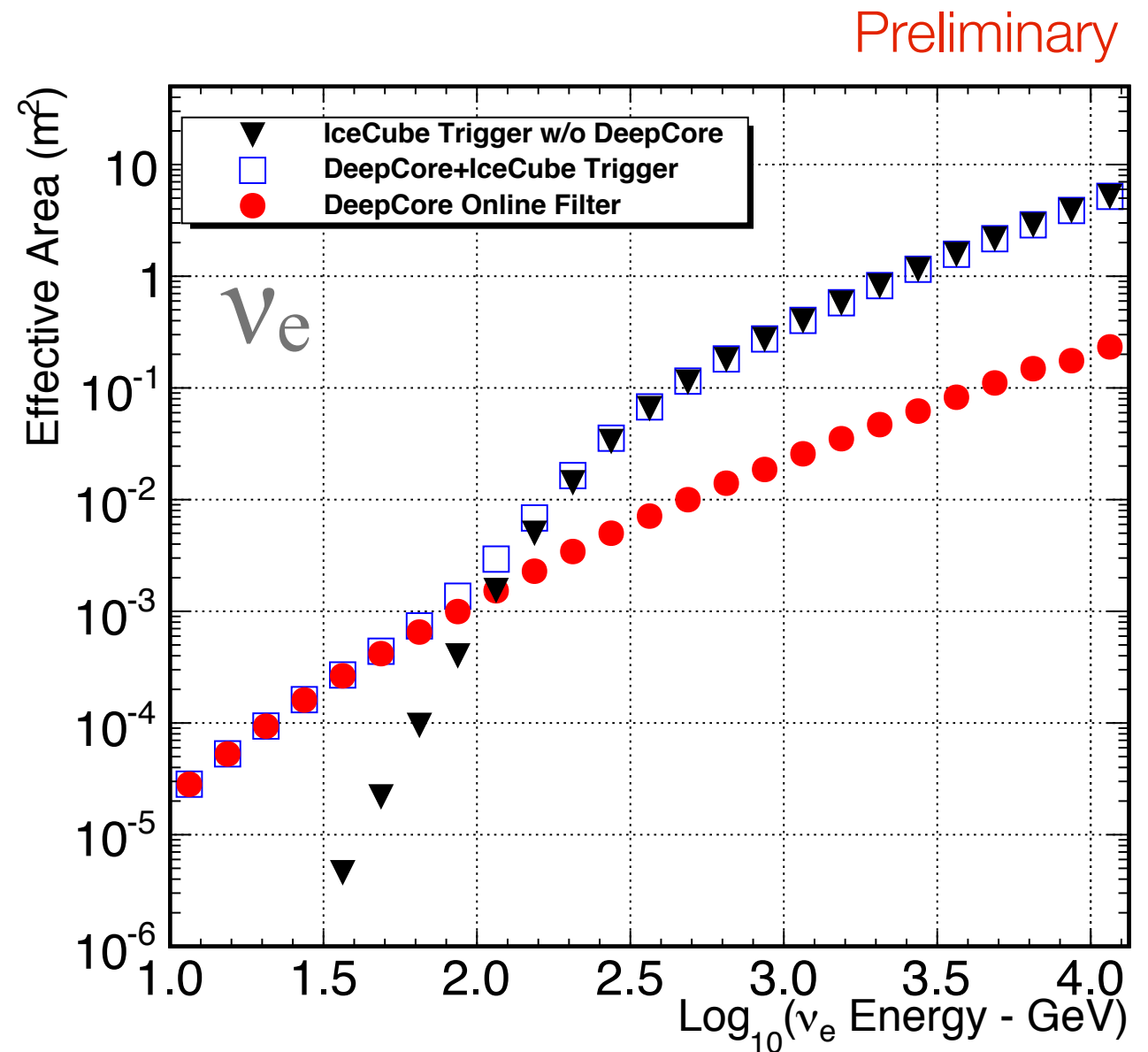
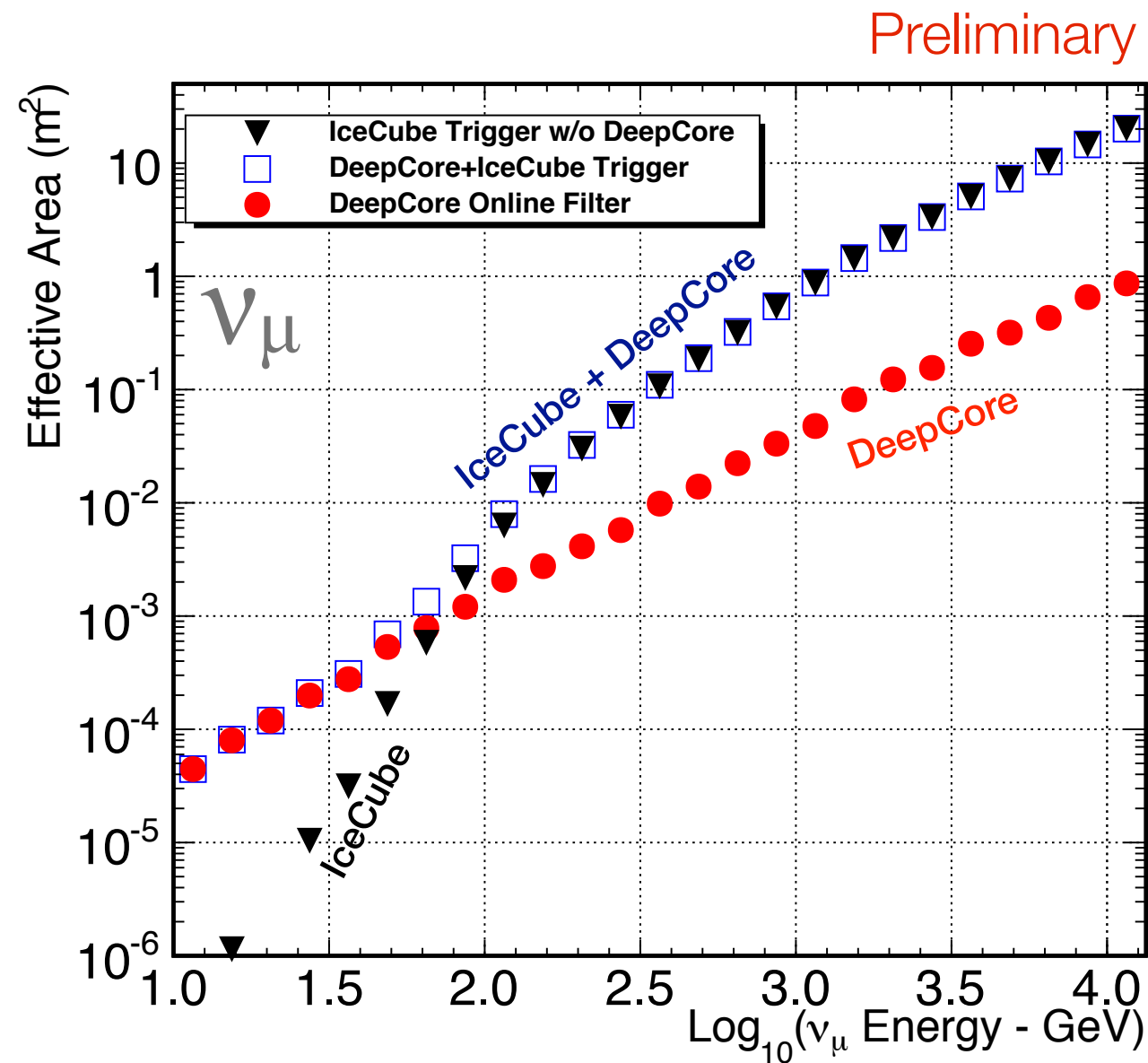
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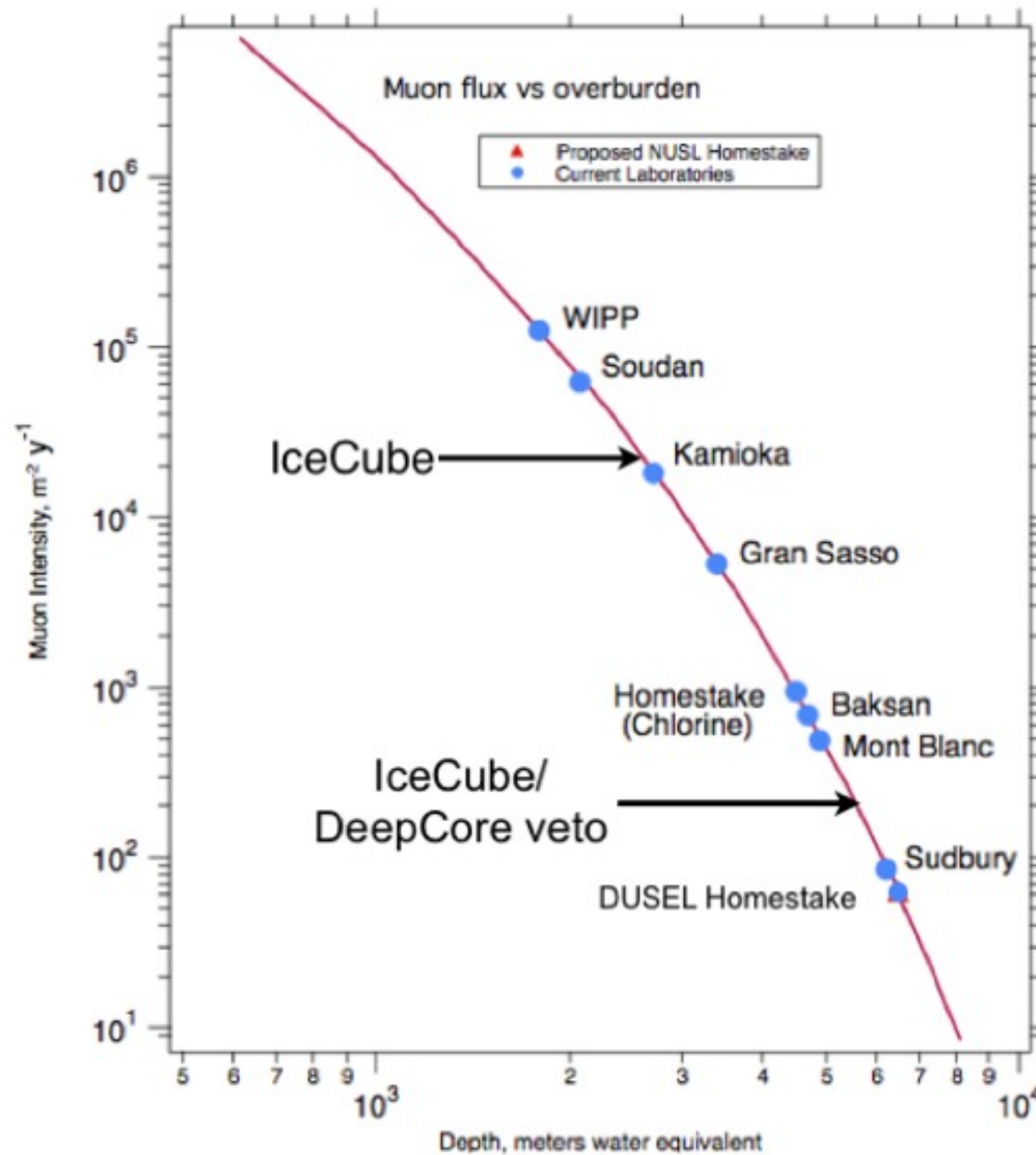
# Thanks

# Backup

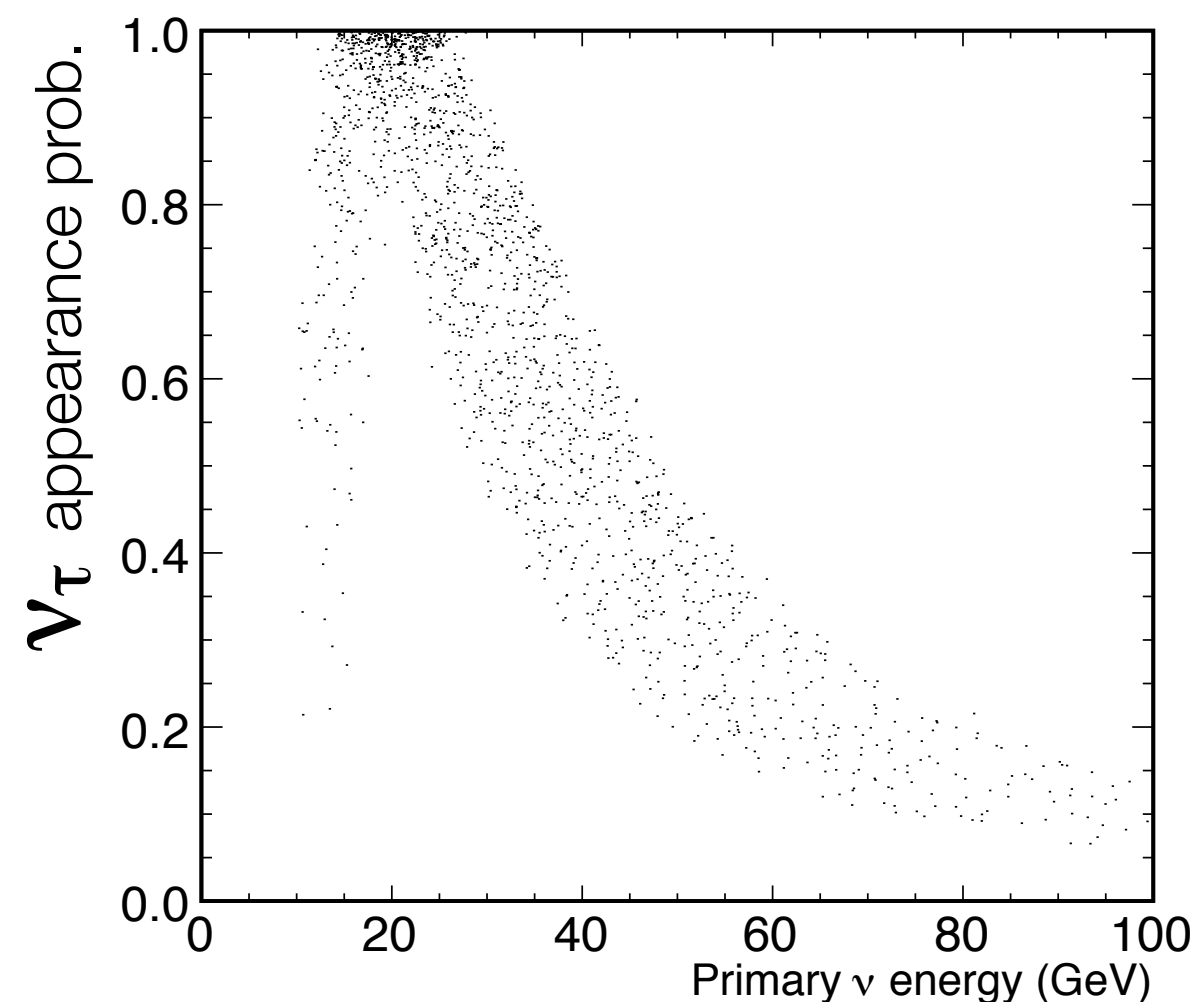
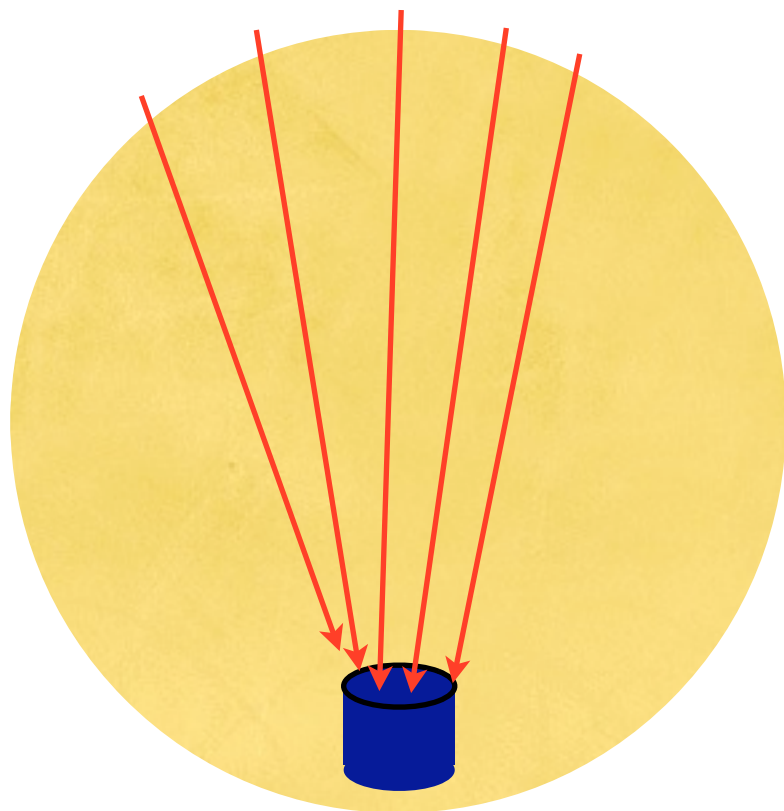
# Effective Neutrino Area

- IceCube
- DeepCore
- Beyond DeepCore



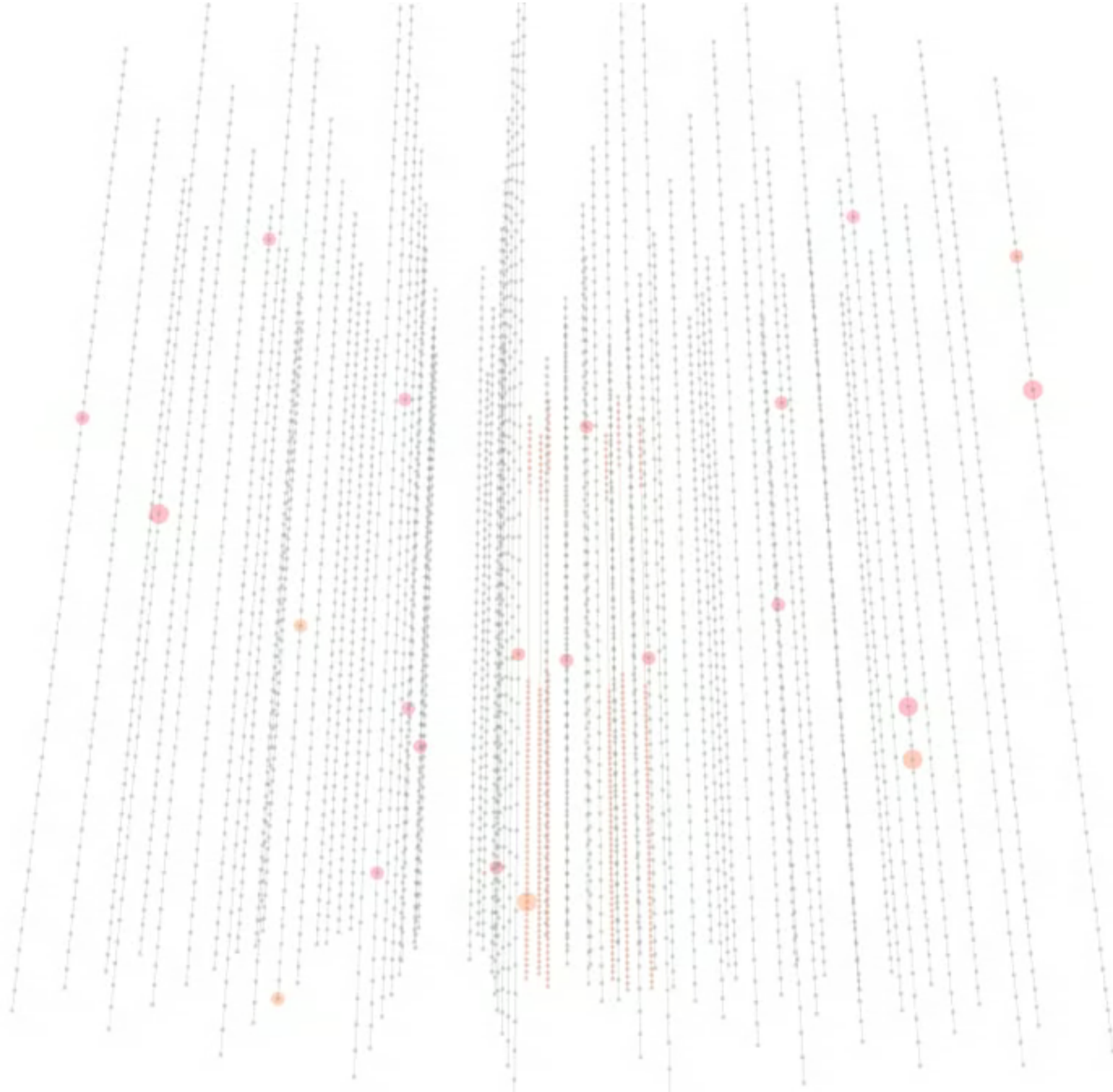


- $\nu_\tau$  events, in addition to neutral current and  $\nu_e$  events, produce cascades
- Signal will be smeared
  - Neutrino-lepton opening angle
  - Zenith angle resolution

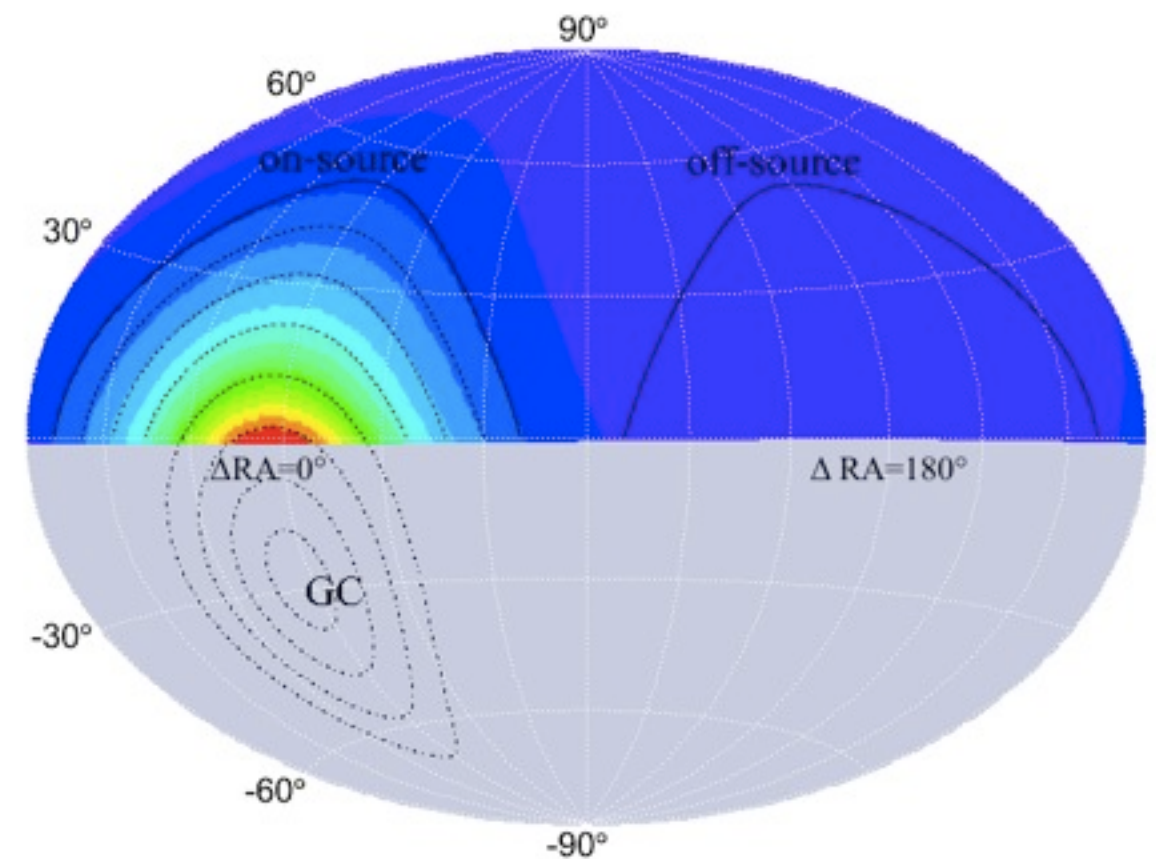


# Neutrino Candidate 2

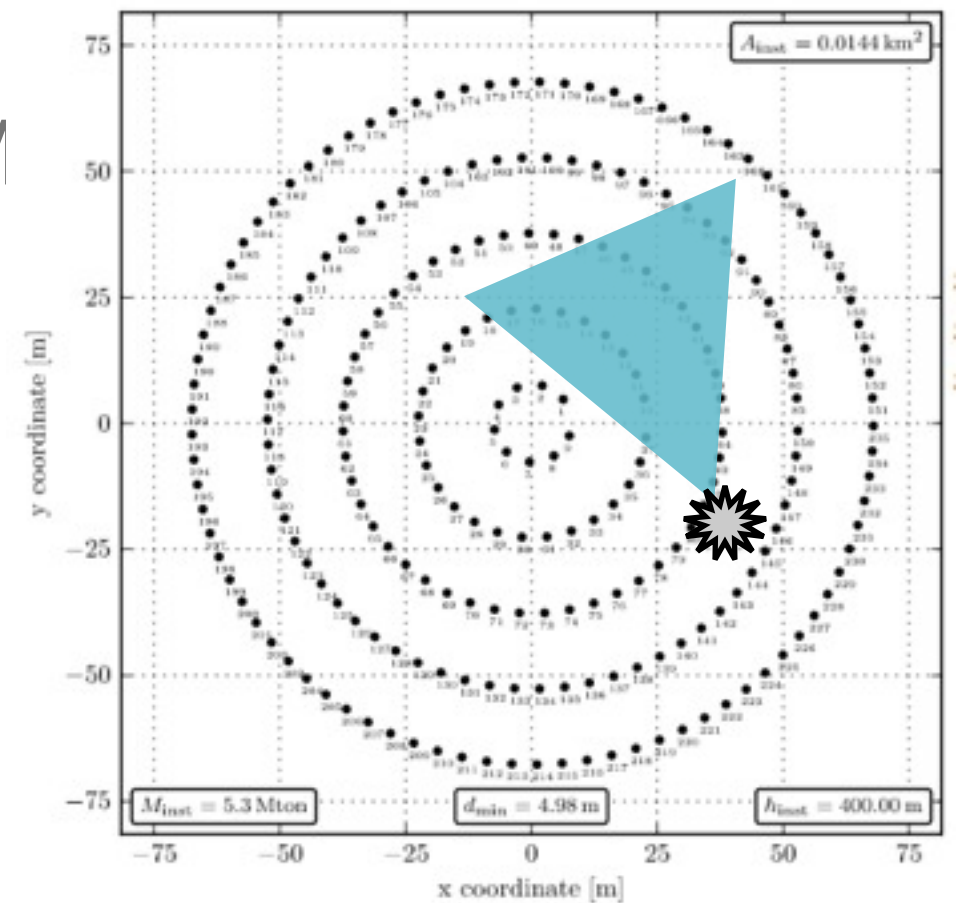
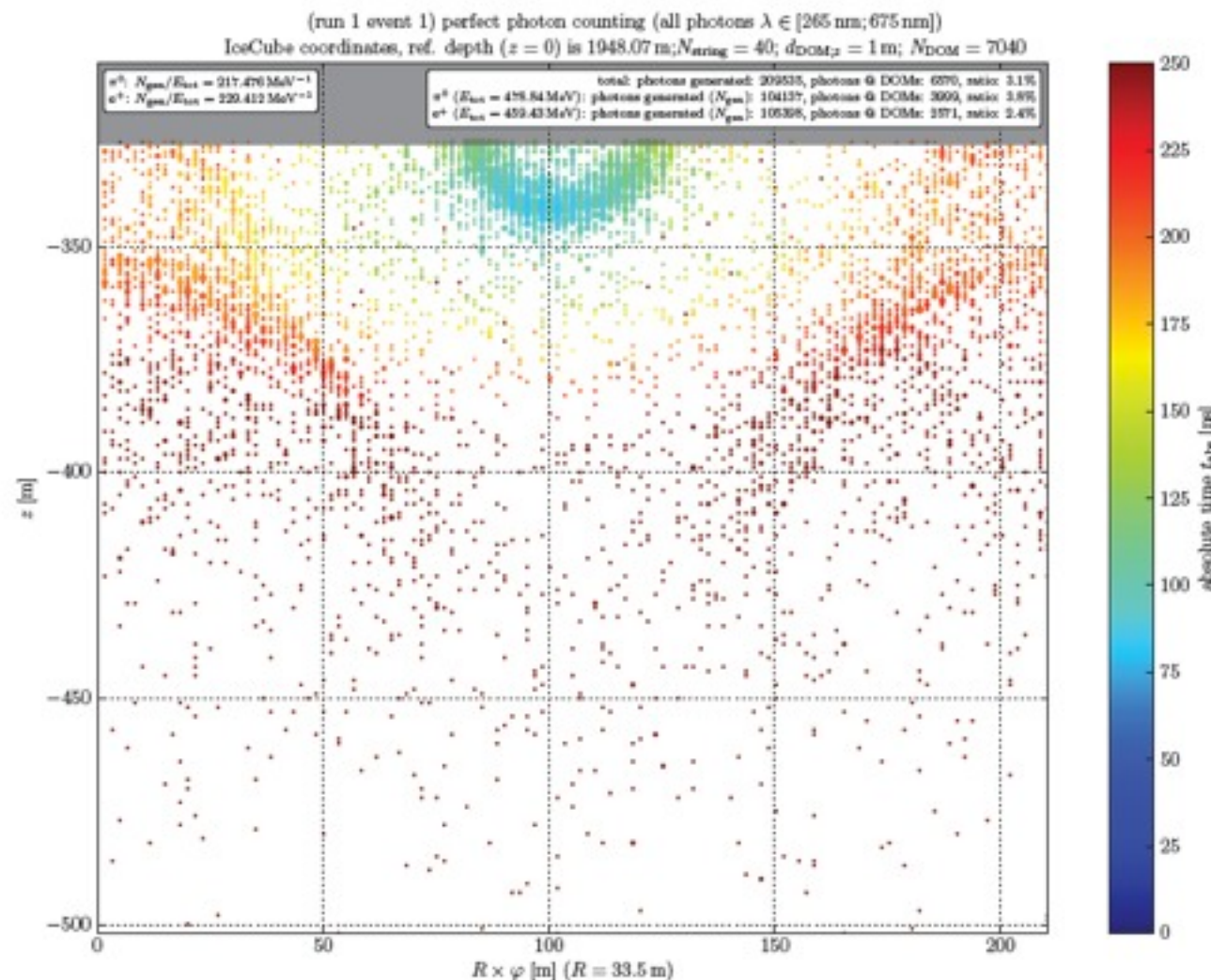
- IceCube
- DeepCore
- Beyond DeepCore



- Signal depends on Halo distribution(DM density) model and SUSY model(DM annihilation channels)
- on-source versus off-source anisotropy
- IC22 analysis did not observe a large scale anisotropy and has a 90% C.L. on WIMP self-annihilation cross-section of  $\langle\sigma_{AV}\rangle 10^{-22} \text{ cm}^3\text{s}^{-1}$  in the WIMP mass range of 200 GeV-10 TeV



- At 100% photon collection efficiency, cylindrical deployment and 1 meter OM spacing, cerenkov ring from proton decay is visible
- Simulation and testing is under active development



## Cost Estimate for a One Megaton Detector

- Costs are driven completely by total photocathode area
  - Is there a more cost-efficient way to collect Cherenkov photons?
- Costs seem competitive, even if management, contingency, personnel, etc. increase the total
- Scaling up to larger volume would be roughly linear in cost
  - Scaling *down* might be harder – how much photocathode can we pack in per unit volume?

